

Наблюдение осциляций электронных антинейтрино в эксперименте Дая Бэй

А.Г.Ольшевский, ОИЯИ

Марковские чтения, 12 мая 2012г.



Neutrino Oscillation

Neutrinos change flavor (e, μ, τ) with time

Principle:

Mass eigenstates \neq Interaction (flavor) eigenstates

$$|\nu_e\rangle = \sum_{m_i} U_{ei}^* |\nu_i\rangle$$

Physical Parameters: (chosen by nature)

θ :

3 angles between mass/flavor eigenstates set **oscillation amplitude**

Δm^2 :

Differences in 3 neutrino masses determine **oscillation frequency** (distance)

We want to know θ and Δm^2

First Evidence of Oscillation:

Davis detects 1/3 expected solar neutrinos (1968)



A Decade of Progress

Clear experimental evidence of neutrino oscillation in recent years

$$U_{if} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \times \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \times \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$c_{ij} \equiv \cos \theta_{ij}$ and $s_{ij} \equiv \sin \theta_{ij}$

$$\theta_{23} \approx 45^\circ$$

Atmospheric v
Accelerator v

$$\theta_{13} < 10^\circ$$

Short-Baseline Reactor v
Accelerator v

$$\theta_{12} \approx 35^\circ$$

Solar v
Long-Baseline Reactor v

θ_{13} : Only angle not yet firmly observed.



11 May 2012



Наблюдение осцилляций электронных нейтрино в эксперименте Дая Бэй", А.Г.Ольшевский, ОИЯИ



Recent Hints of θ_{13}

2011 has given many hints:

Solar + KamLAND: G.L.Fogli *et al.*, Phys. Rev. D 84, 053007 (2011)

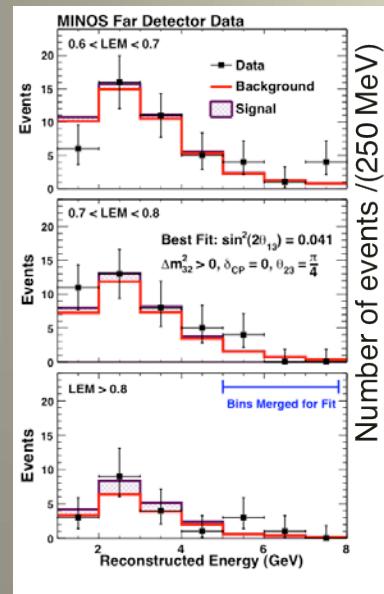
MINOS: P. Adamson *et al.*, Phys. Rev. Lett. 107, 181802 (2011)

T2K: K. Abe *et al.*, Phys. Rev. Lett. 107 041801 (2011)

Double CHOOZ: Y. Abe *et al.*, arXiv:1112.6353

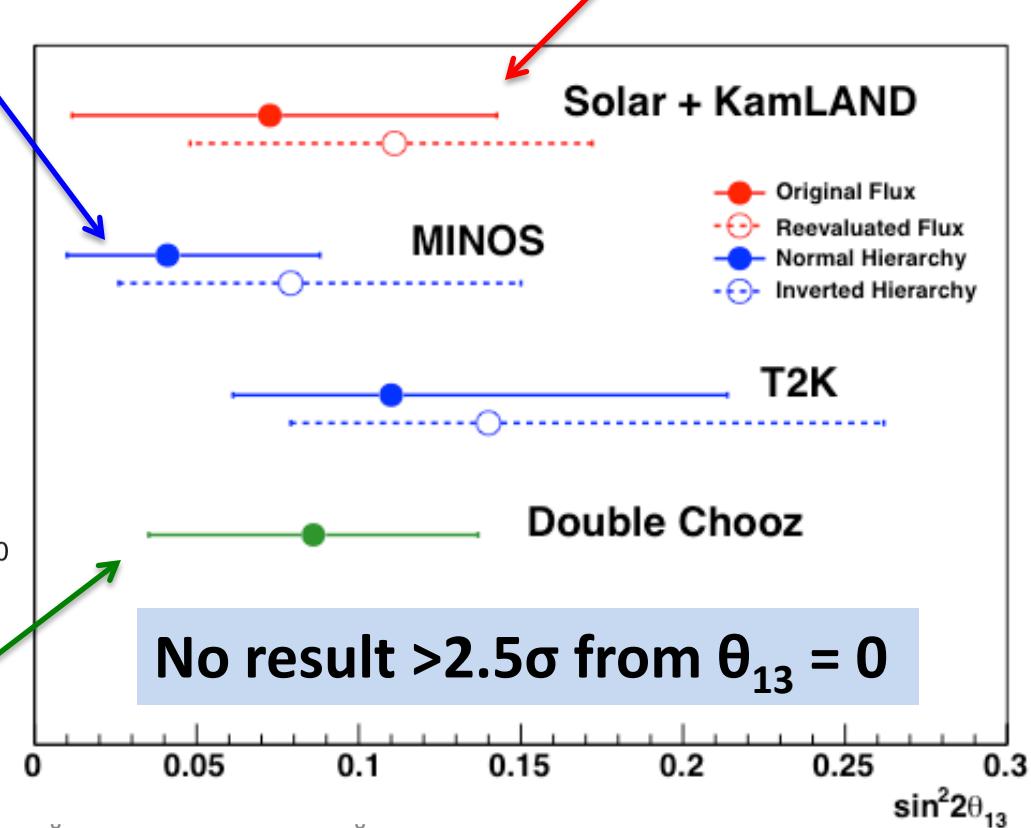
Tension between solar, reactor ν_e oscillations suggest $\theta_{13} > 0$.

Appearance of ν_e in ν_μ accelerator beam



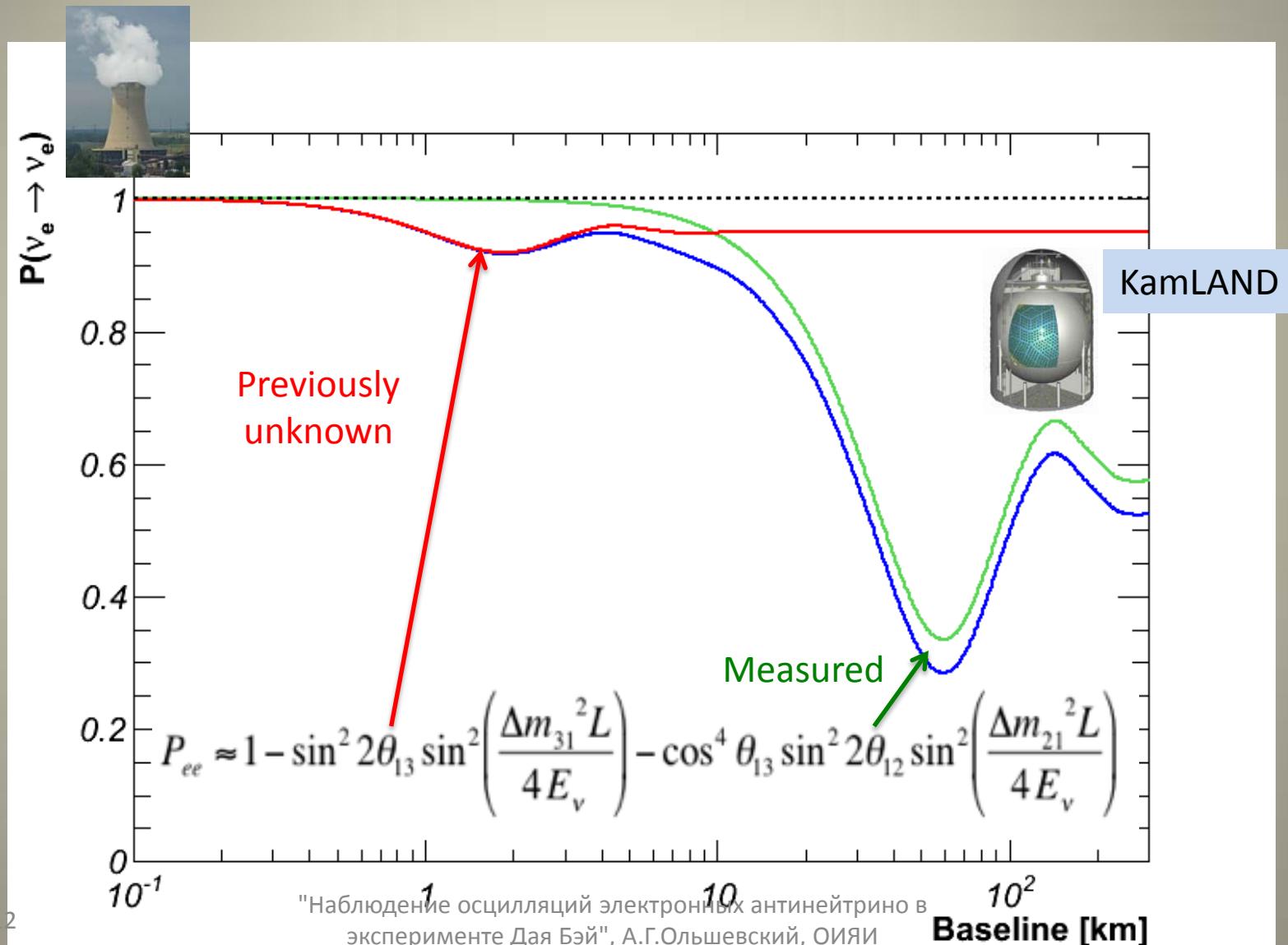
Double Chooz reported improved single detector measurement.

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Reactor Neutrino Oscillation

θ_{13} revealed by a deficit of reactor antineutrinos at ~ 2 km.



Relative Measurement

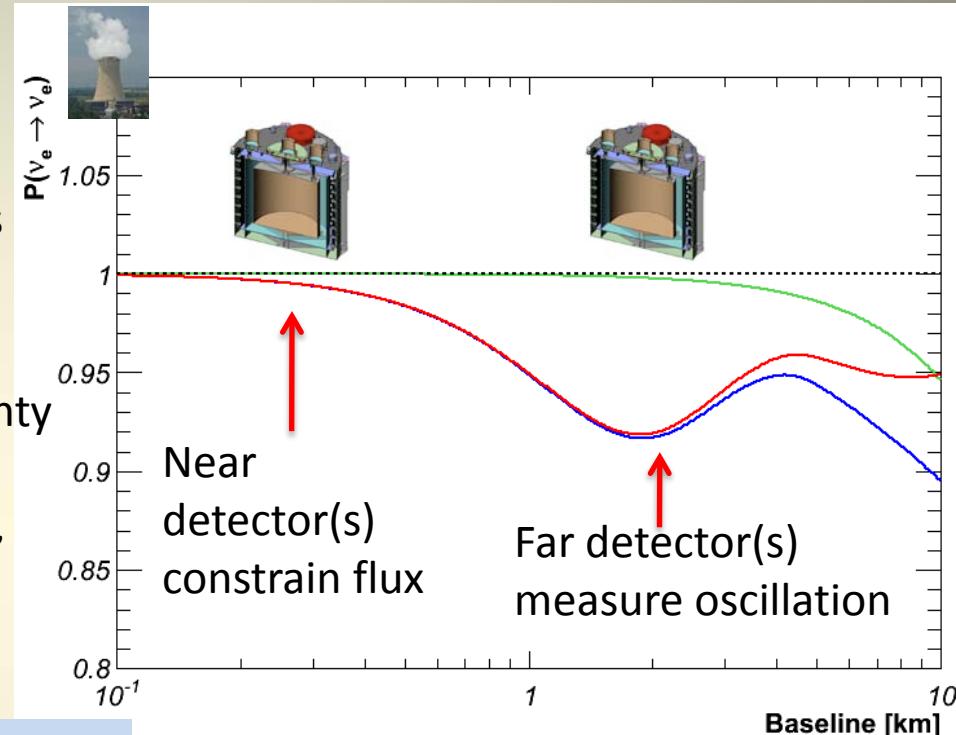
Absolute Reactor Flux:

Largest uncertainty in previous measurements

Relative Measurement:

Multiple detectors removes absolute uncertainty

First proposed by L. A. Mikaelyan and V.V. Sinev,
Phys. Atomic Nucl. 63 1002 (2000)



Far/Near ν_e Ratio

Distances from reactor

Oscillation deficit

$$\frac{N_f}{N_n} = \left(\frac{N_{p,f}}{N_{p,n}} \right) \left(\frac{L_n}{L_f} \right)^2 \left(\frac{\epsilon_f}{\epsilon_n} \right) \left[\frac{P_{\text{sur}}(E, L_f)}{P_{\text{sur}}(E, L_n)} \right]$$

Detector Target Mass

Detector efficiency

Изучение электронных антинейтрино в
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Daya Bay: An Ideal Location

17.4 GW (thermal) reactor power adjacent to mountains.

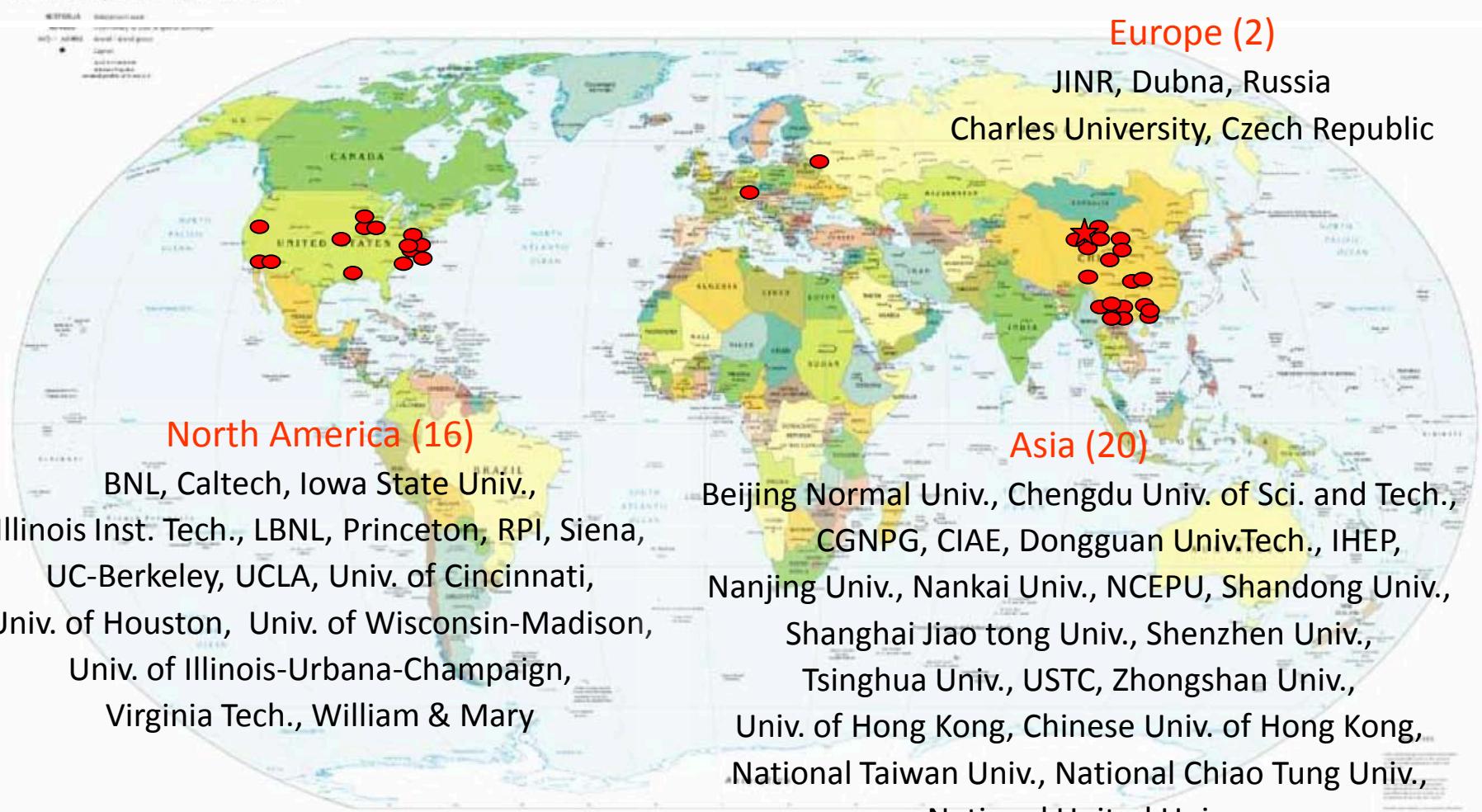


Reactors produce $\sim 2 \times 10^{20}$ antineutrinos / s / GW

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The Daya Bay Collaboration

Political Map of the World, June 1999



~230 Collaborators

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Experiment Layout

**6 Antineutrino Detectors (ADs)
in 3 underground halls.**

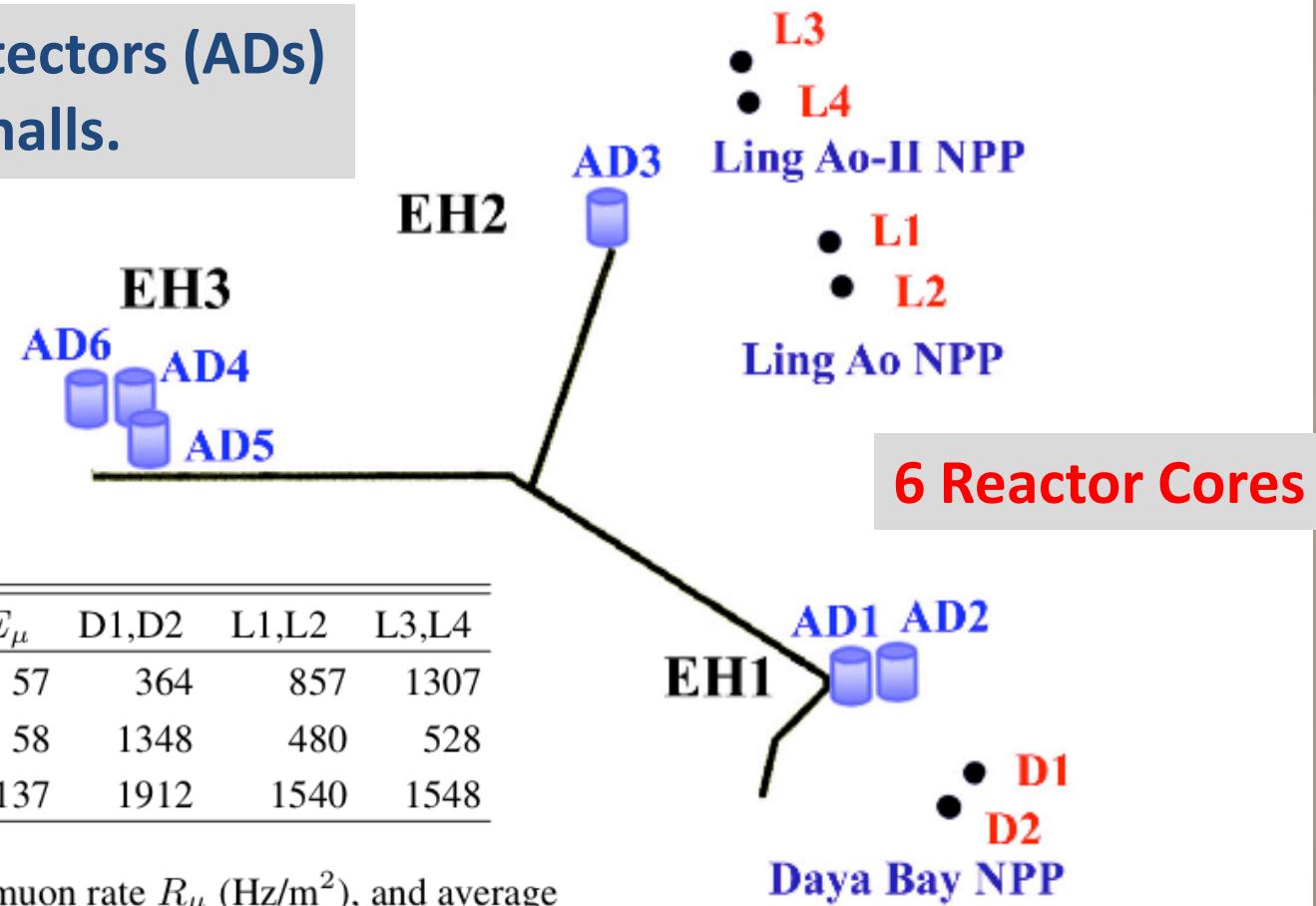


TABLE I. Overburden (m.w.e), muon rate R_μ (Hz/m²), and average muon energy E_μ (GeV) of the three EHs, and the distances (m) to the reactor pairs.

Experiment Survey

Negligible reactor flux uncertainty (<0.02%) from precise survey.

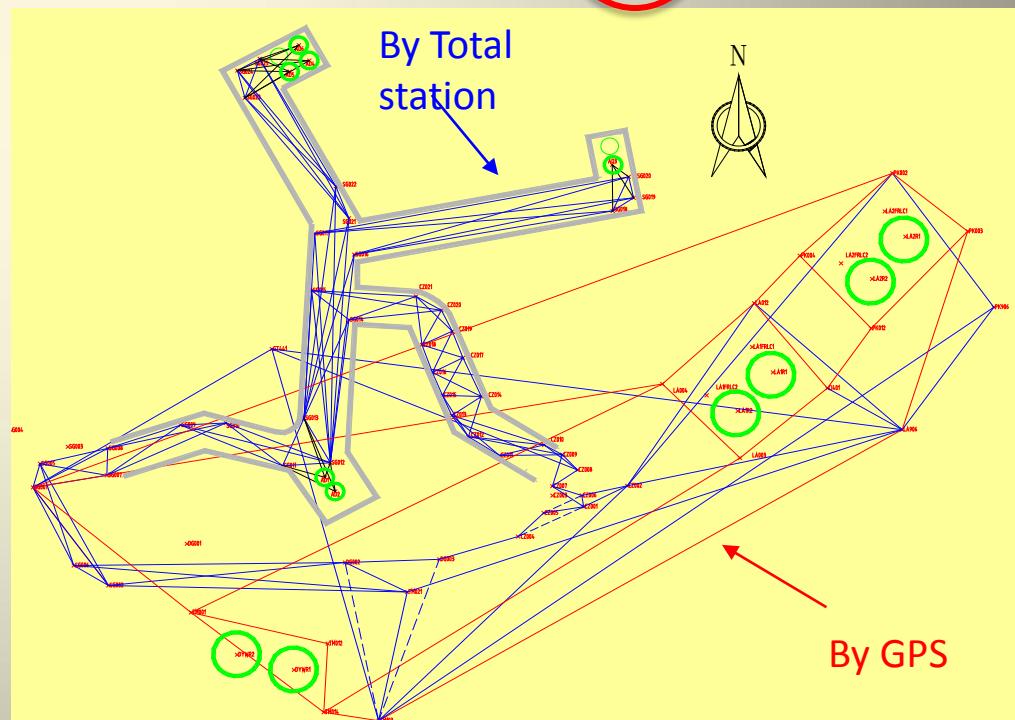
Detailed Survey:

- GPS above ground
- Total Station underground
- Final precision: 28mm

$$\frac{N_f}{N_n} = \left(\frac{N_{p,f}}{N_{p,n}} \right) \left(\frac{L_n}{L_f} \right)^2 \left(\frac{\epsilon_f}{\epsilon_n} \right) \left[\frac{P_{\text{sur}}(E, L_f)}{P_{\text{sur}}(E, L_n)} \right]$$

Validation:

- Three independent calculations
- Cross-check survey
- Consistent with reactor plant and design plans



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Antineutrino Detectors

6 ‘functionally identical’ detectors:

Reduce systematic uncertainties

$$\frac{N_f}{N_n} = \left(\frac{N_{p,f}}{N_{p,n}} \right) \left(\frac{L_n}{L_f} \right)^2 \left(\frac{\epsilon_f}{\epsilon_n} \right) \left[\frac{P_{\text{sur}}(E, L_f)}{P_{\text{sur}}(E, L_n)} \right]$$

3 nested cylinders:

Inner: 20 tons Gd-doped LS ($d=3.1\text{m}$)

Mid: 20 tons LS ($d=4\text{m}$)

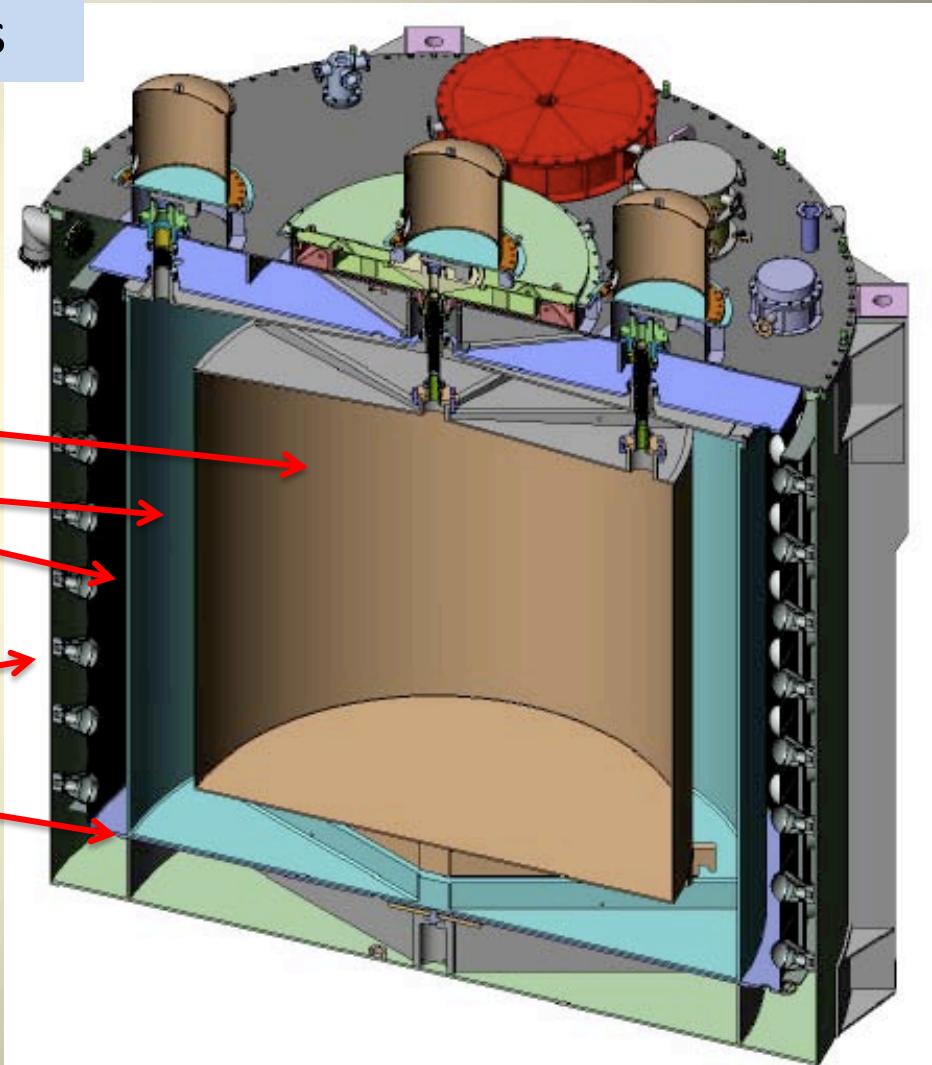
Outer: 40 tons mineral oil buffer ($d=5\text{m}$)

Each detector:

192 8-inch Photomultipliers

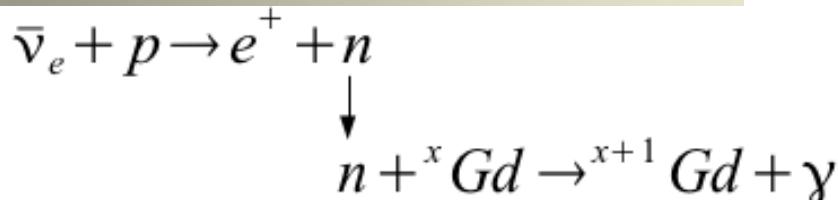
Reflectors at top/bottom of cylinder

Provides $(7.5 / \sqrt{E} + 0.9)\%$ energy resolution



Detection Method

Inverse β-decay (IBD):



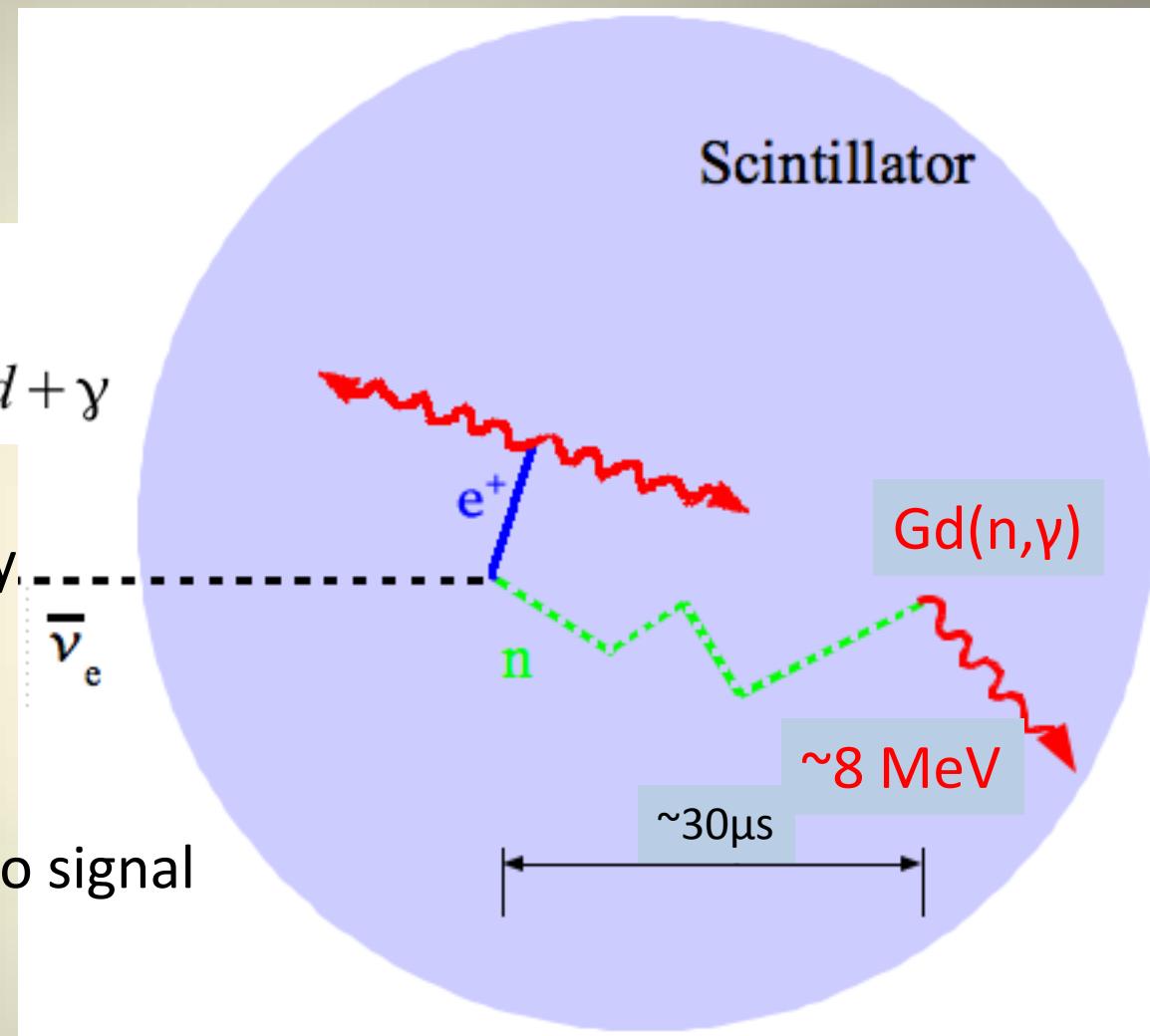
Prompt positron:

Carries antineutrino energy

$$E_{e^+} \approx E_\nu - 0.8 \text{ MeV}$$

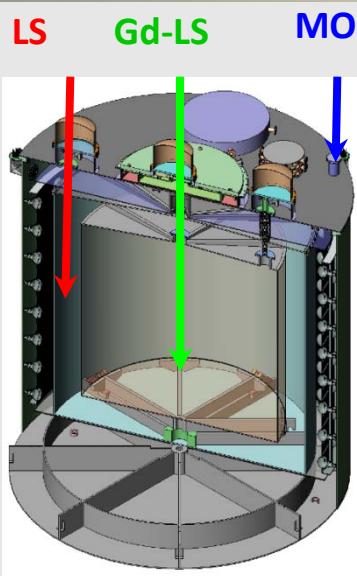
Delayed neutron capture:

Efficiently tags antineutrino signal



Prompt + Delayed coincidence provides distinctive signature

Detector Filling



Detector target filled from GdLS in ISO tank.

Load cells measure
20 ton target mass
to 3 kg (0.015%)

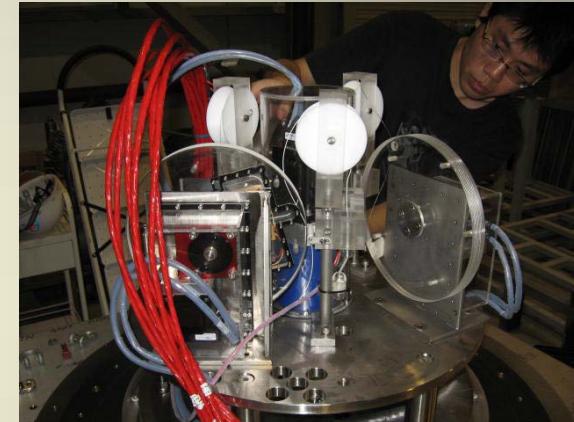
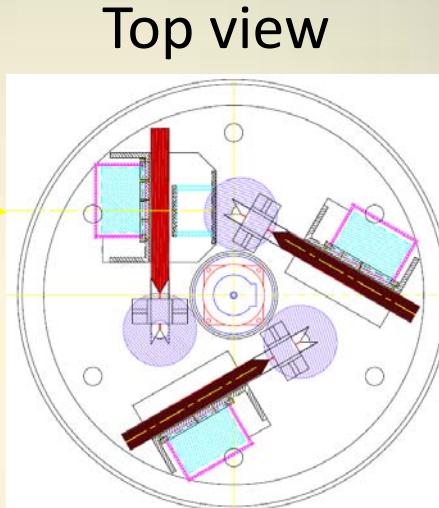
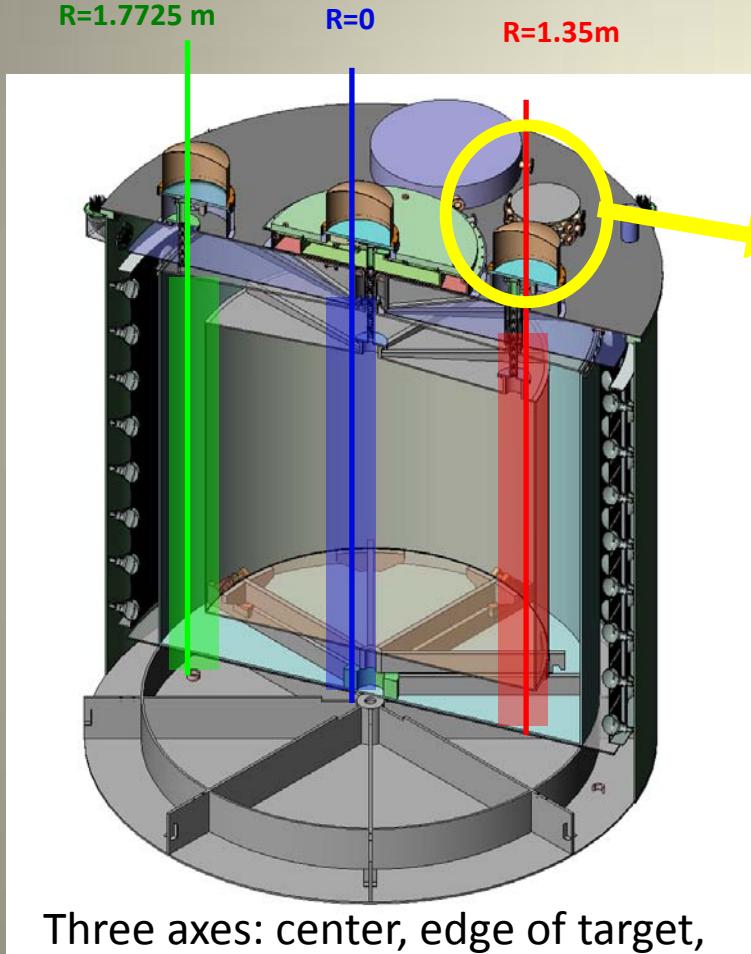
$$\frac{N_f}{N_n} = \left(\frac{N_{p,f}}{N_{p,n}} \right) \left(\frac{L_n}{L_f} \right)^2 \left(\frac{\epsilon_f}{\epsilon_n} \right) \left[\frac{P_{\text{sur}}(E, L_f)}{P_{\text{sur}}(E, L_n)} \right]$$

- 3 fluids filled simultaneously, with heights matched to minimize stress on acrylic vessels**
- **Gadolinium-doped Liquid Scintillator (GdLS)**
 - **Liquid Scintillator (LS)**
 - **Mineral Oil (MO)**

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Automated Calibration System

3 Automatic calibration ‘robots’ (ACUs) on each detector

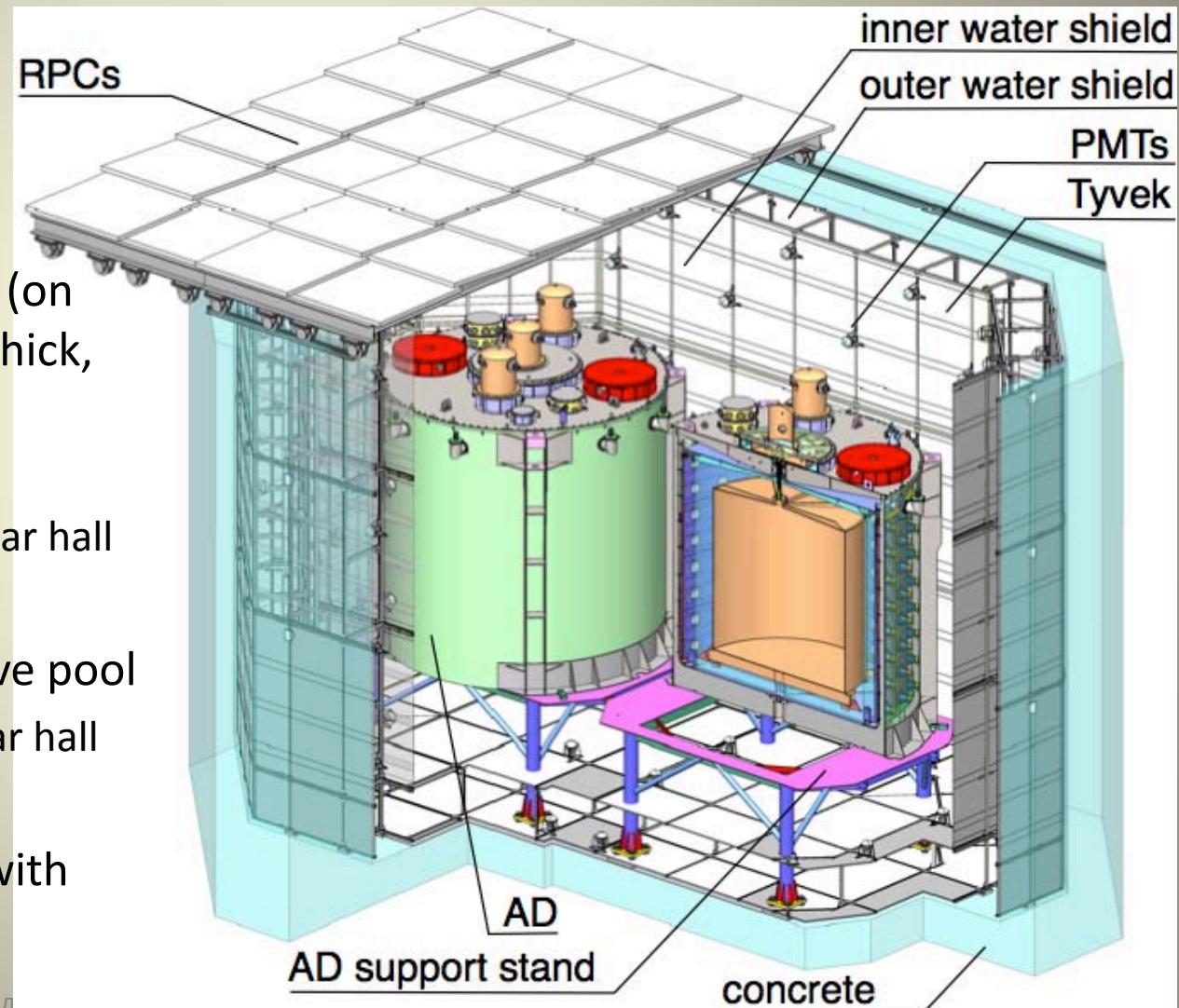


3 sources for each z axis on a turntable (position accuracy < 5 mm):

- 10 Hz ^{68}Ge ($0 \text{ KE } e^+ = 2 \times 0.511 \text{ MeV } \gamma$'s)
- 0.5 Hz ^{241}Am - ^{13}C neutron source ($3.5 \text{ MeV } n$ without γ) + 100 Hz ^{60}Co gamma source ($1.173 + 1.332 \text{ MeV } \gamma$)
- LED diffuser ball (500 Hz) for T_0 and gain

Muon Tagging System

Dual tagging systems: 2.5 meter thick two-section water shield and RPCs



- Outer layer of water veto (on sides and bottom) is 1m thick, inner layer >1.5m. Water extends 2.5m above ADs
 - 288 8" PMTs in each near hall
 - 384 8" PMTs in Far Hall
- 4-layer RPC modules above pool
 - 54 modules in each near hall
 - 81 modules in Far Hall
- Goal efficiency: > 99.5% with uncertainty <0.25%

JINR contribution to Daya Bay

- Liquid Scintillator measurements and optimization:**

- Light Yield
- Transparency
- Energy Resolution
- Neutron capture for Gd loaded LS



Scintillator layer height, cm	Total detection efficiency of < 0,4 eV thermal neutrons, %	
	Gd-LS	LS
1	12 ± 2	–
2	19 ± 3	–
3,5	29 ± 4	–
4,5	35 ± 5	17 ± 3



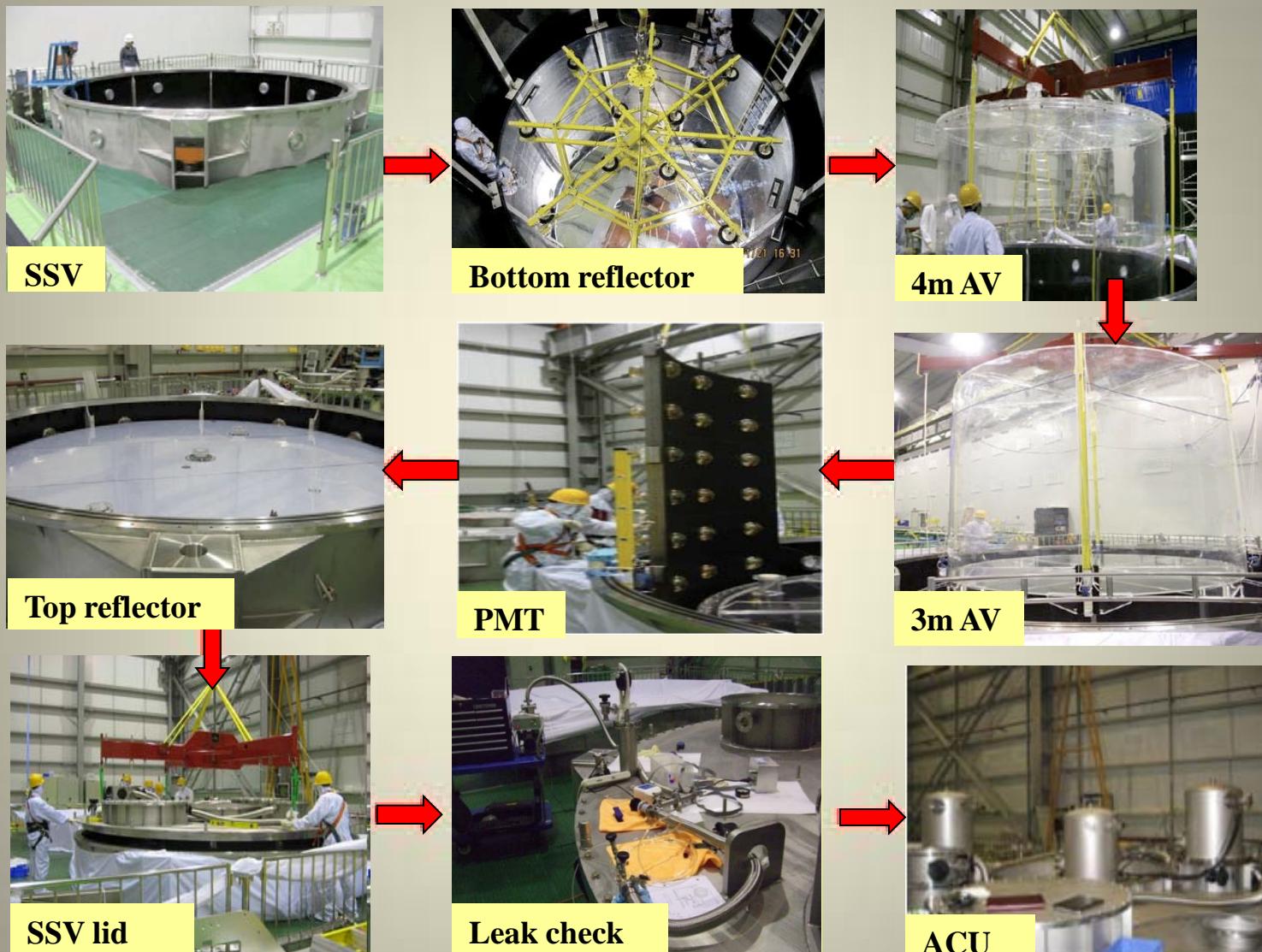
- Technology of PPO production was restored in the JINR Member State Ukraine and 1.5t of PPO were produced and delivered to Daya Bay**

- Data analysis:**

- Background simulation
- Oscillation Analysis

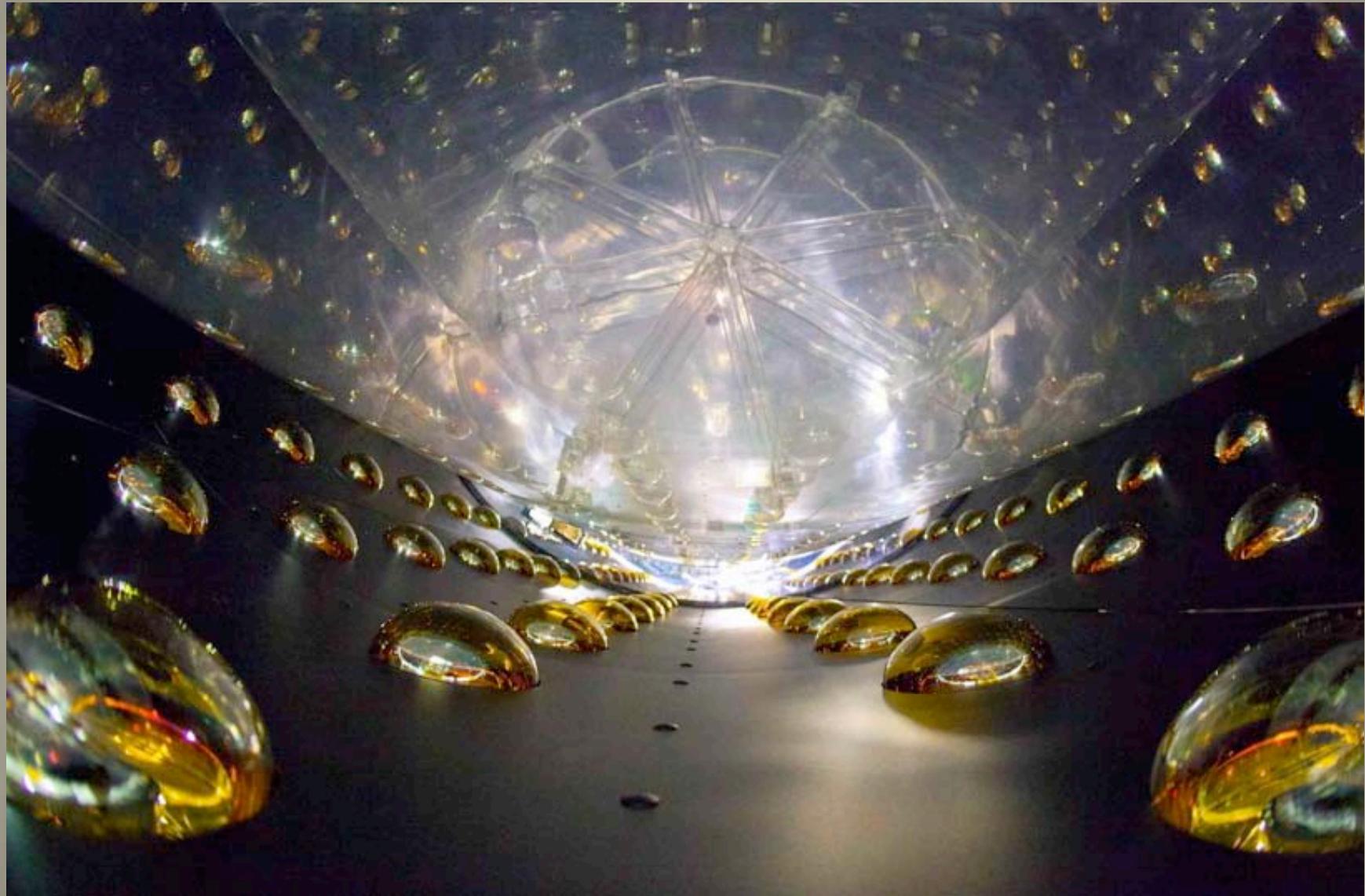
"Наблюдение осцилляций электронных антинейтрино в эксперименте Дая Бэй", А.Г.Ольшевский, ОИЯИ

Detector Assembly



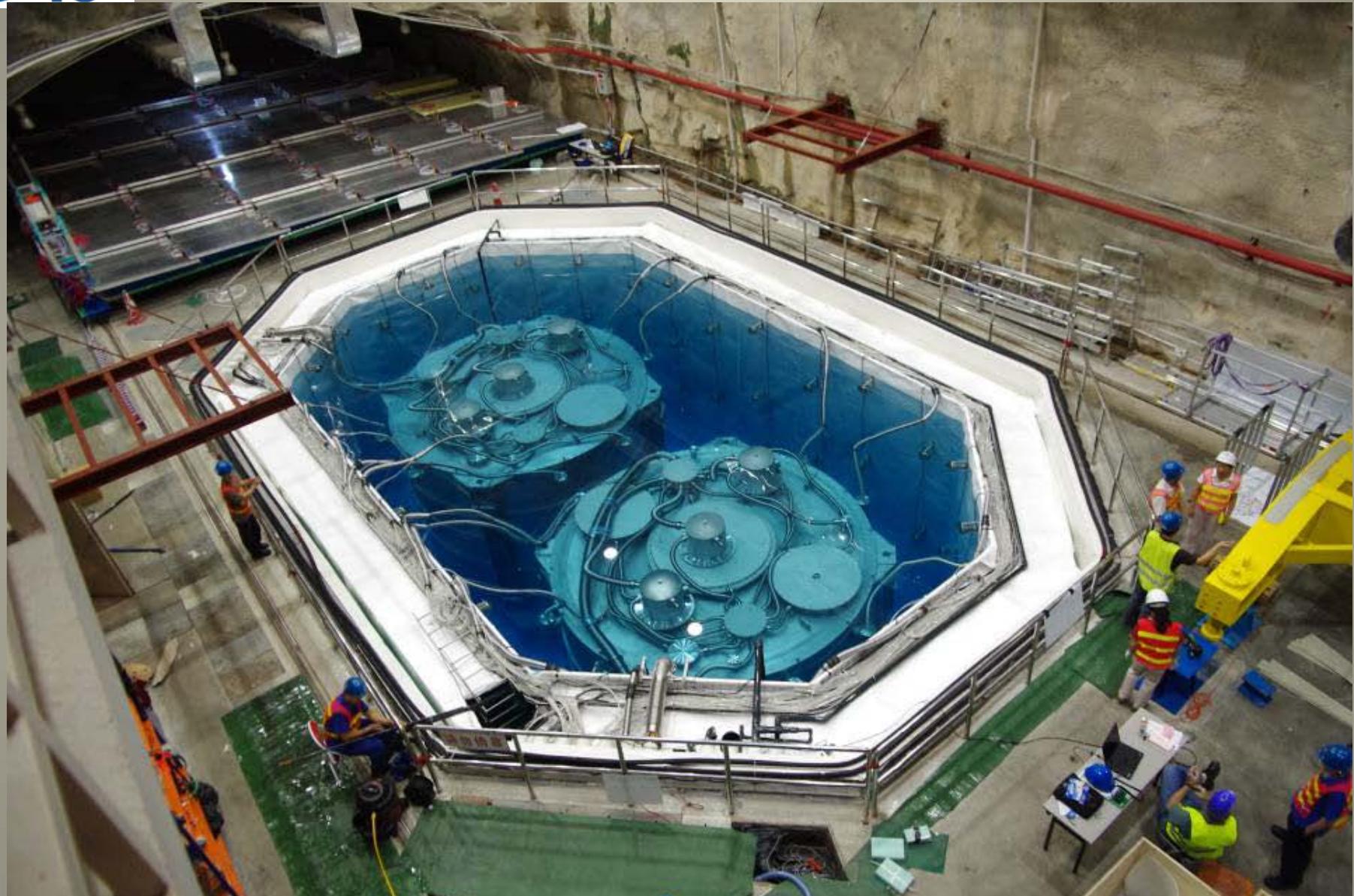
"Наблюдение осцилляций электронных антинейтрино в эксперименте Дая Бэй", А.Г.Ольшевский, ОИЯИ

Interior of Antineutrino Detector



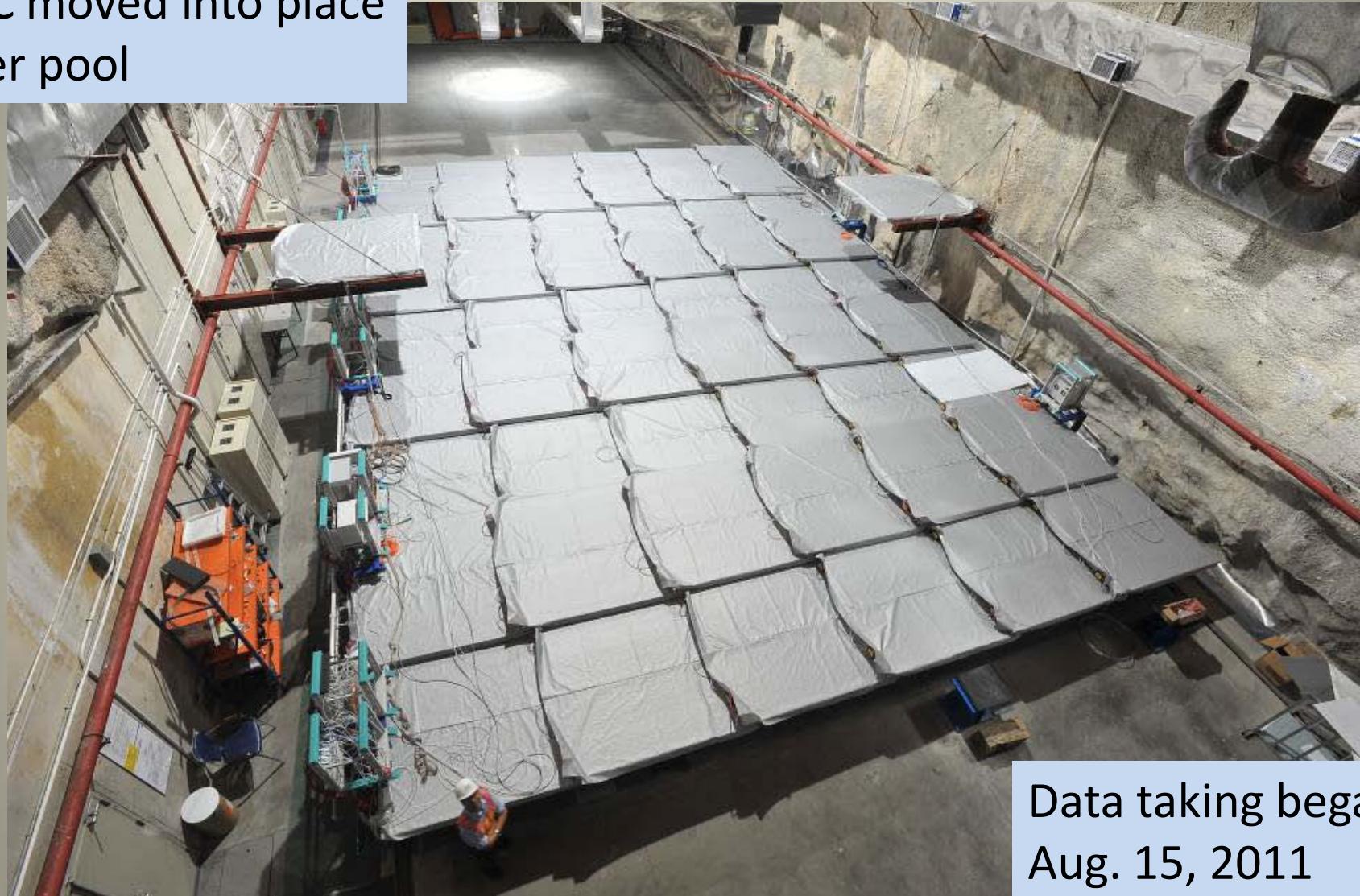
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EH1: Pool Filled



Hall 1: Completed

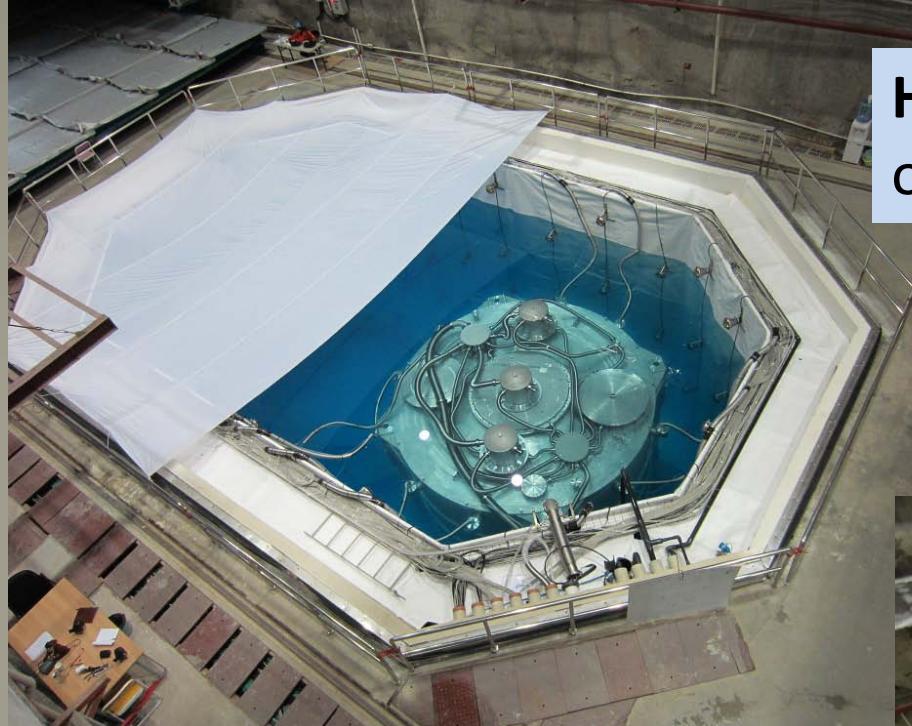
RPC moved into place over pool



Data taking began
Aug. 15, 2011

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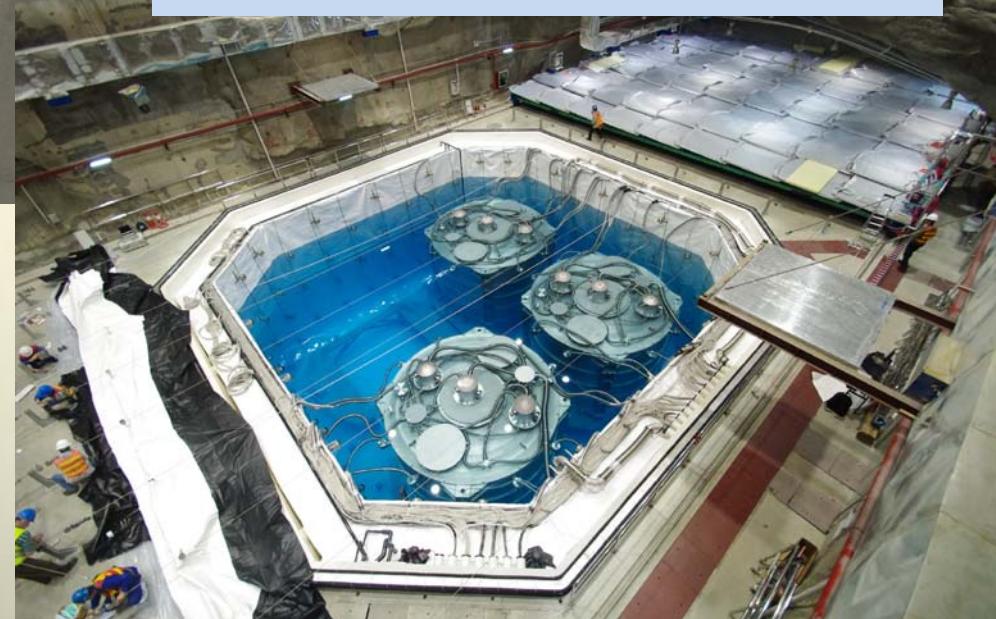
Hall 2 and Hall 3



Hall 2: Began 1 AD operation
on Nov. 5, 2011



Hall 3: Began 3 AD operation
on Dec. 24, 2011



2 more ADs still in assembly;
installation planned for
Summer 2012

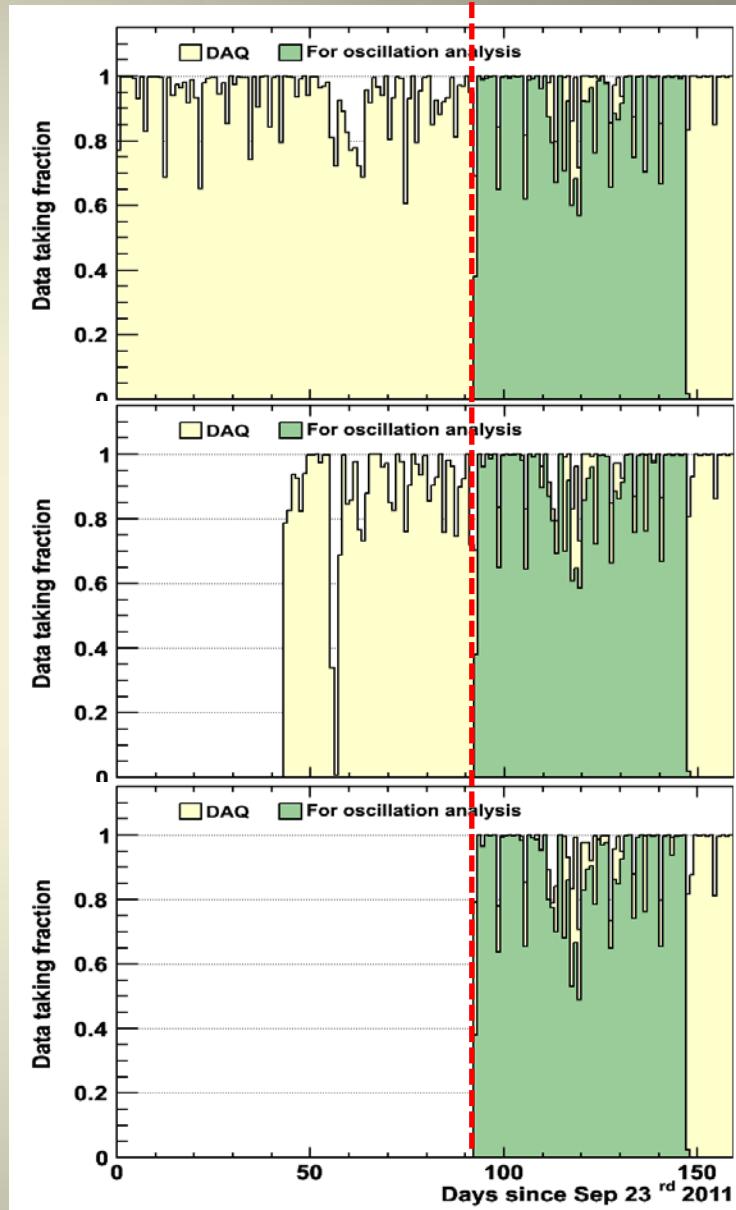
Data Period

Two Detector Comparison:

- Sep. 23, 2011 – Dec. 23, 2011
- Side-by-side comparison of 2 detectors
- Demonstrated detector systematics better than requirements.
- Details presented in:
F.P. An et al., arXiv:1202.6181 (2012)

Current Oscillation Analysis:

- Dec. 24, 2011 – Feb. 17, 2012
- All 3 halls (6 ADs) operating
- DAQ uptime: >97%
- Antineutrino data: ~89%



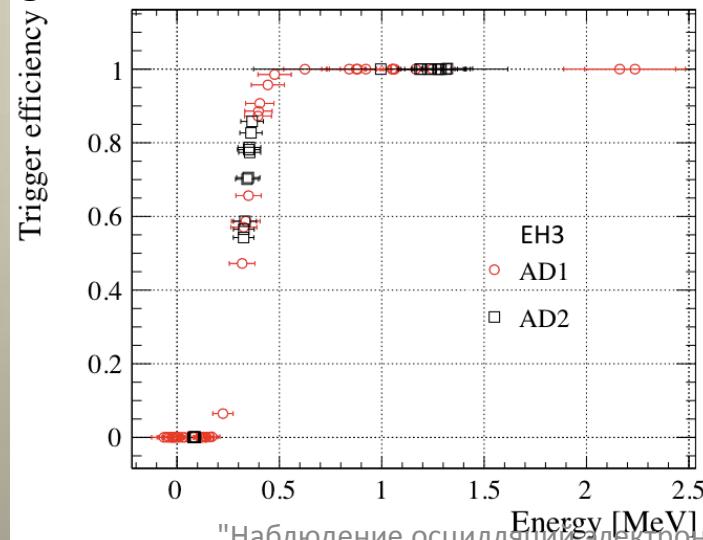
Trigger Performance

Trigger Thresholds:

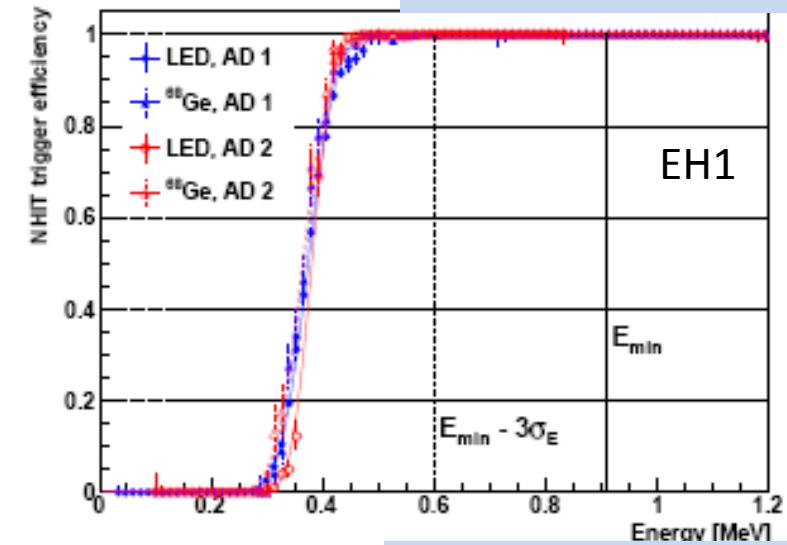
- AD: >45 PMTs (digital trigger)
 >0.4 MeV (analog trigger)
- Inner Water Veto: > 6 PMTs
- Outer Water Veto: >7 PMTs
- RPC: $\frac{3}{4}$ layers in module

Trigger Efficiency:

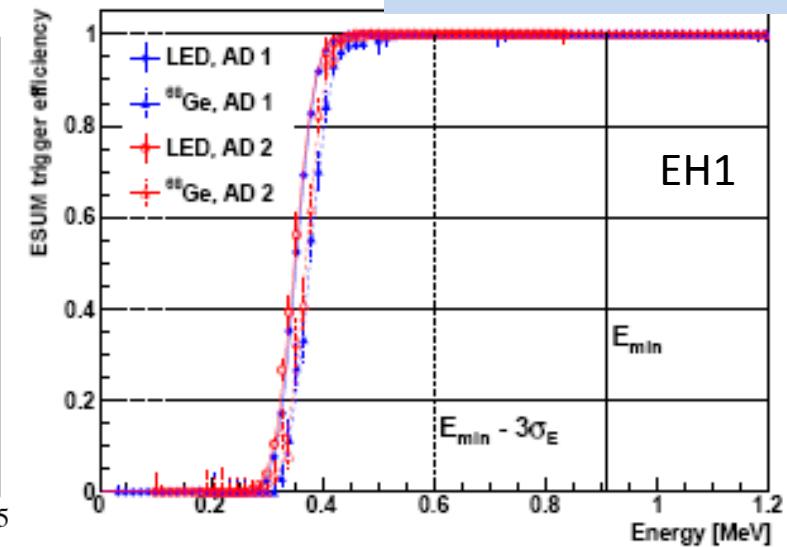
- No measurable inefficiency >0.7 MeV
- Minimum energy expected for prompt antineutrino signal is \sim 0.9 MeV.



Digital Trigger (Nhit)



Analog Trigger (Esum)

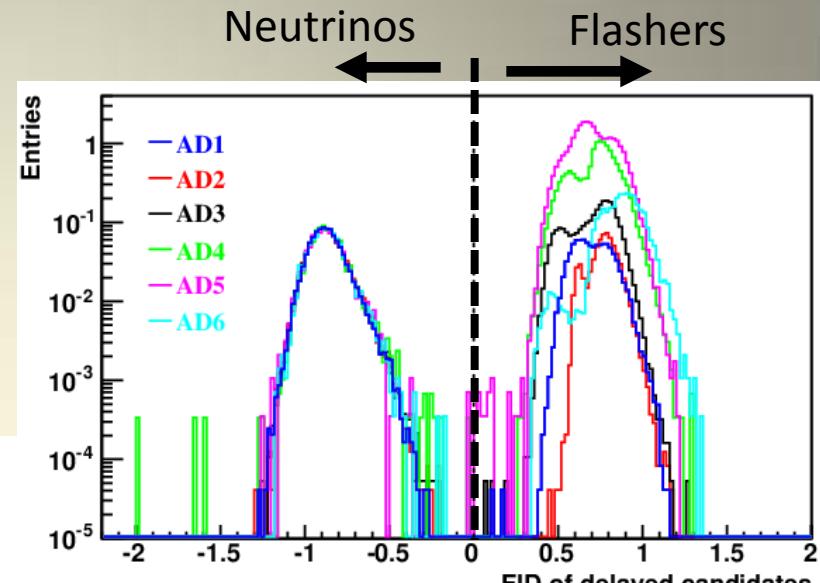
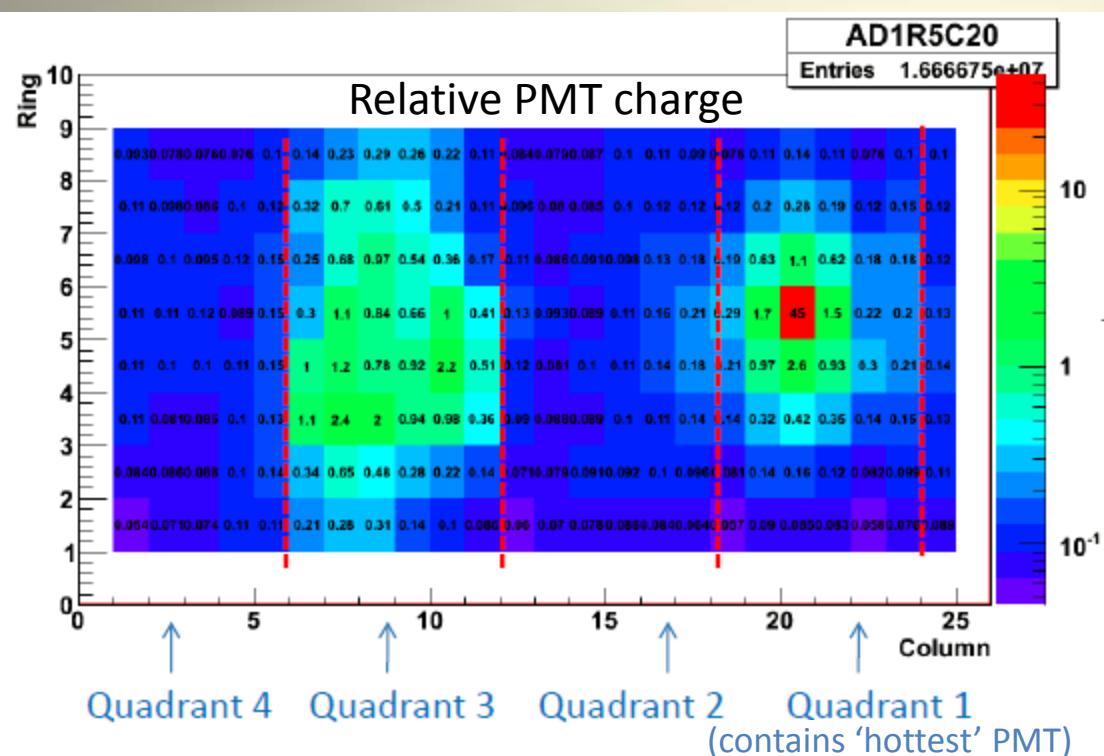


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PMT Light Emission (Flashing)

Flashing PMTs:

- Instrumental background from ~5% of PMTS
- ‘Shines’ light to opposite side of detector
- Easily discriminated from normal signals



$$FID = \log_{10}((MaxQ)^2/(0.45)^2 + (Quad)^2),$$

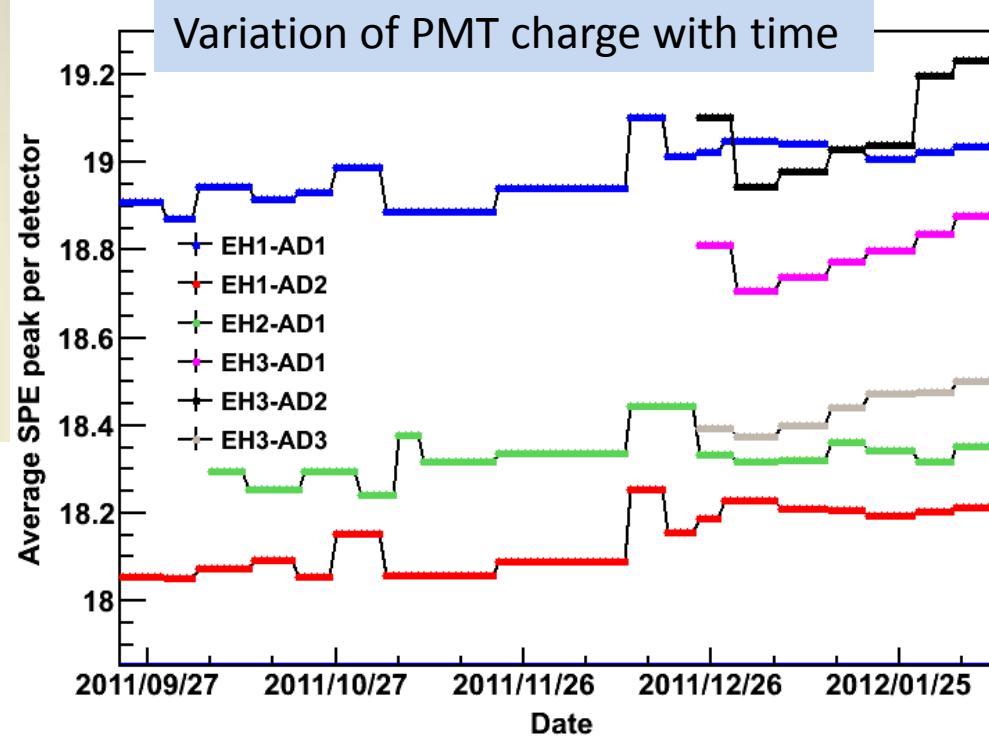
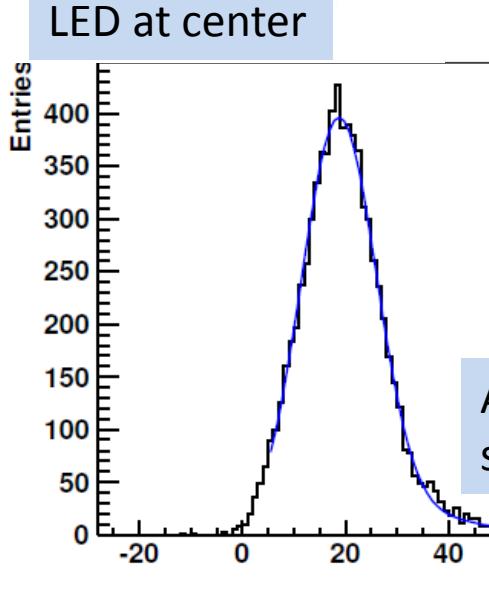
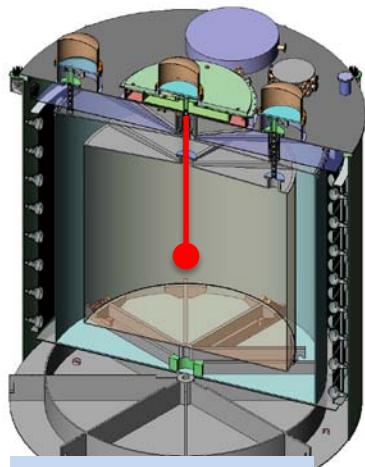
$$Quadrant = Q3/(Q2+Q4)$$

$$MaxQ = maxQ/sumQ$$

Inefficiency to antineutrinos signal:
 $0.024\% \pm 0.006\%(\text{stat})$
 Contamination: < 0.01%

Calibration: PMT Gain

Weekly LED deployments measure charge due to single photons

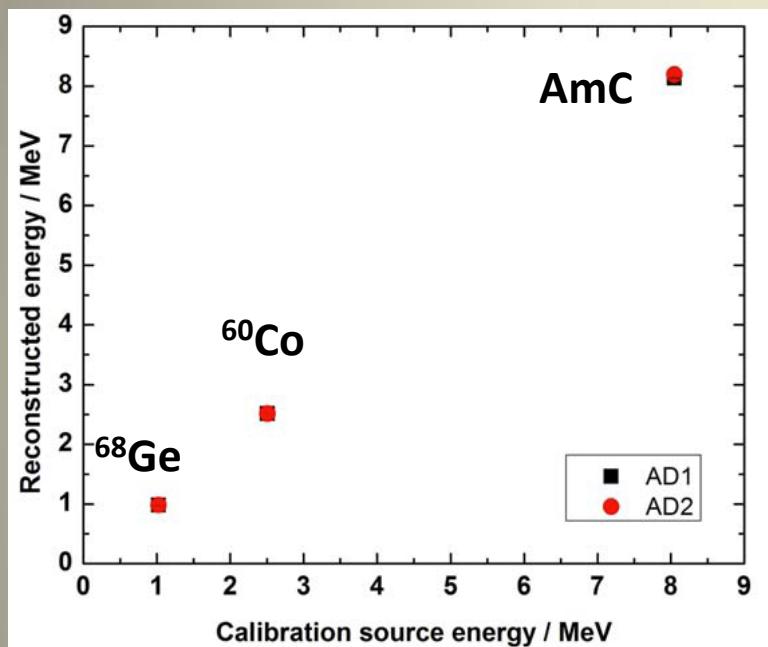


Calibration driven by uncertainty in relative detector efficiency

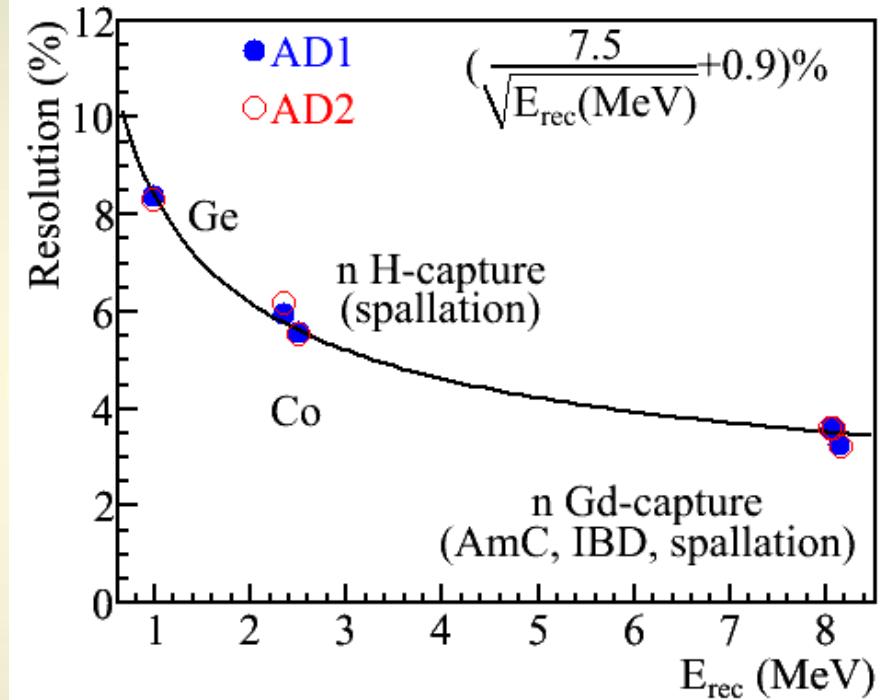
$$\frac{N_f}{N_n} = \left(\frac{N_{p,f}}{N_{p,n}} \right) \left(\frac{L_n}{L_f} \right)^2 \left(\frac{\epsilon_f}{\epsilon_n} \right) \left[\frac{P_{\text{sur}}(E, L_f)}{P_{\text{sur}}(E, L_n)} \right]$$

Calibration: Energy Scale

Energy scale



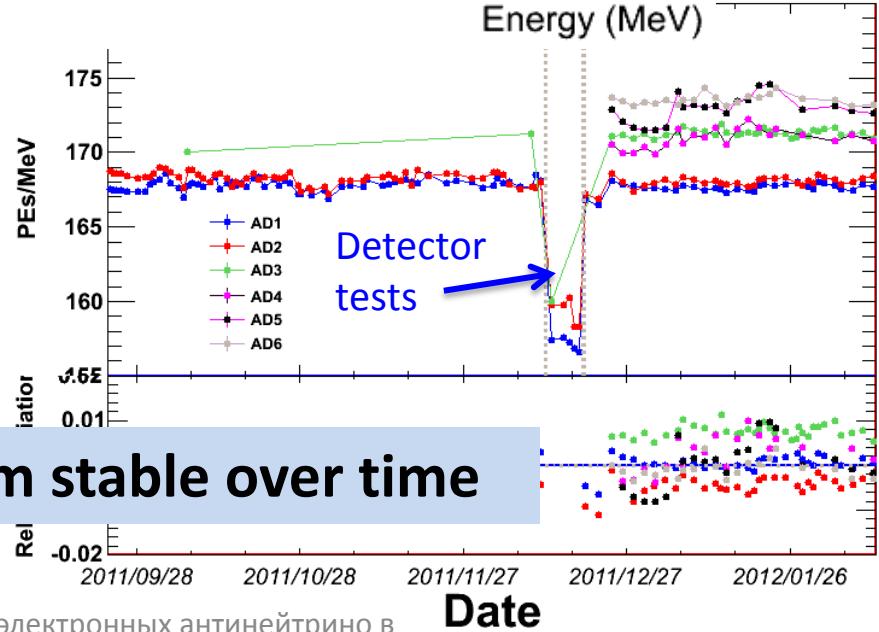
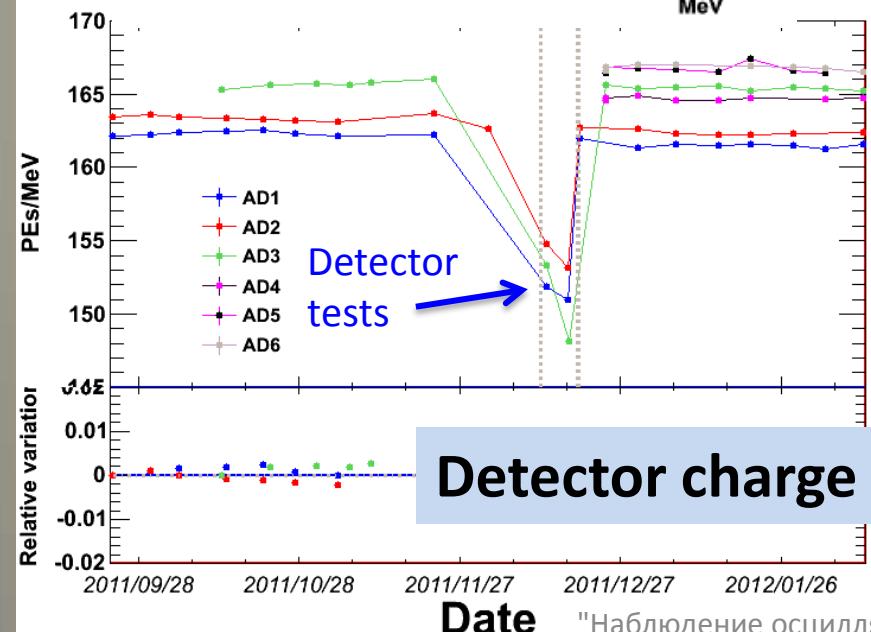
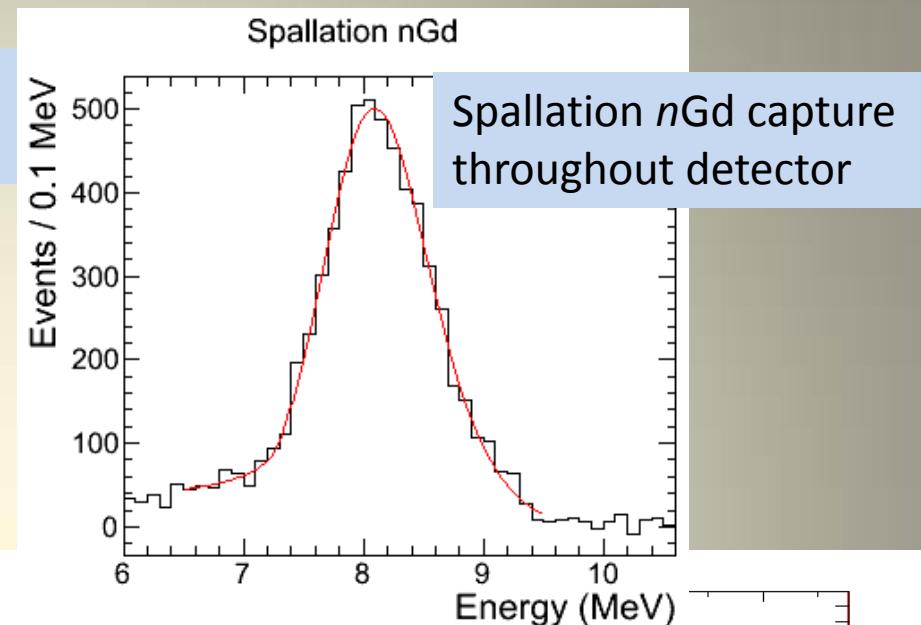
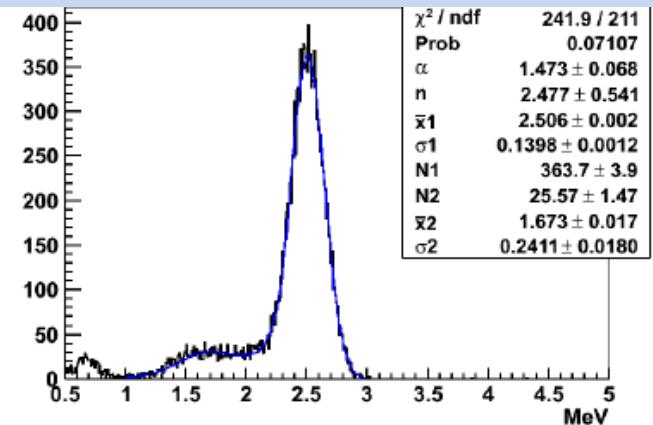
Energy resolution



- Energy linearity at detector centre with ^{68}Ge , ^{60}Co , AmC neutron
- Photon electron yield: $\sim 170 \text{ PE/MeV}$
- Resolution ($\text{RMS}/E_{\text{MeV}}$): $\sim 7.5\%/\sqrt{E_{\text{MeV}}} + 0.9\%$

Calibration: Energy Scale

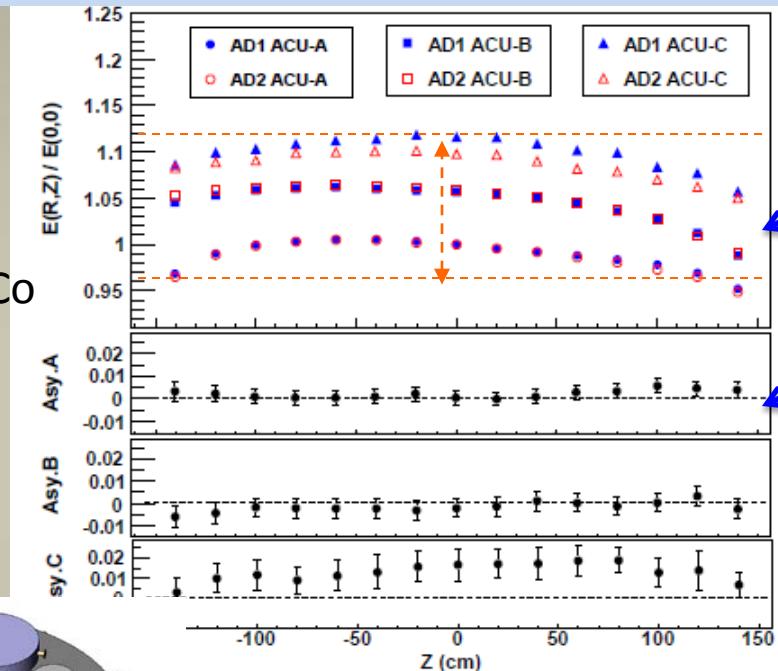
Weekly deployments of ^{60}Co at detector center:
Monitor photoelectrons collected per MeV



Calibration: Detector Uniformity

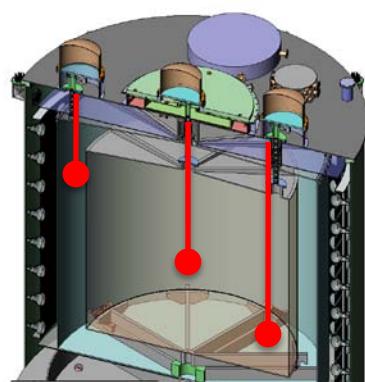
Multiple sources placed along three axes measure uniformity

Example: ^{60}Co



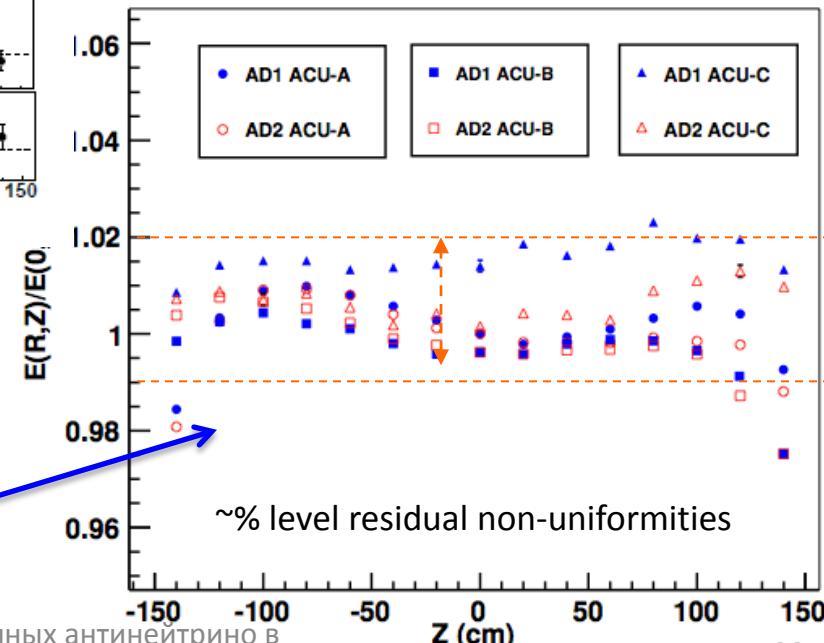
Energy response varies across detector...

...but still consistent between detectors



3 sources along 3 axes

After first-order correction, energy is more uniform.



Singles Spectrum

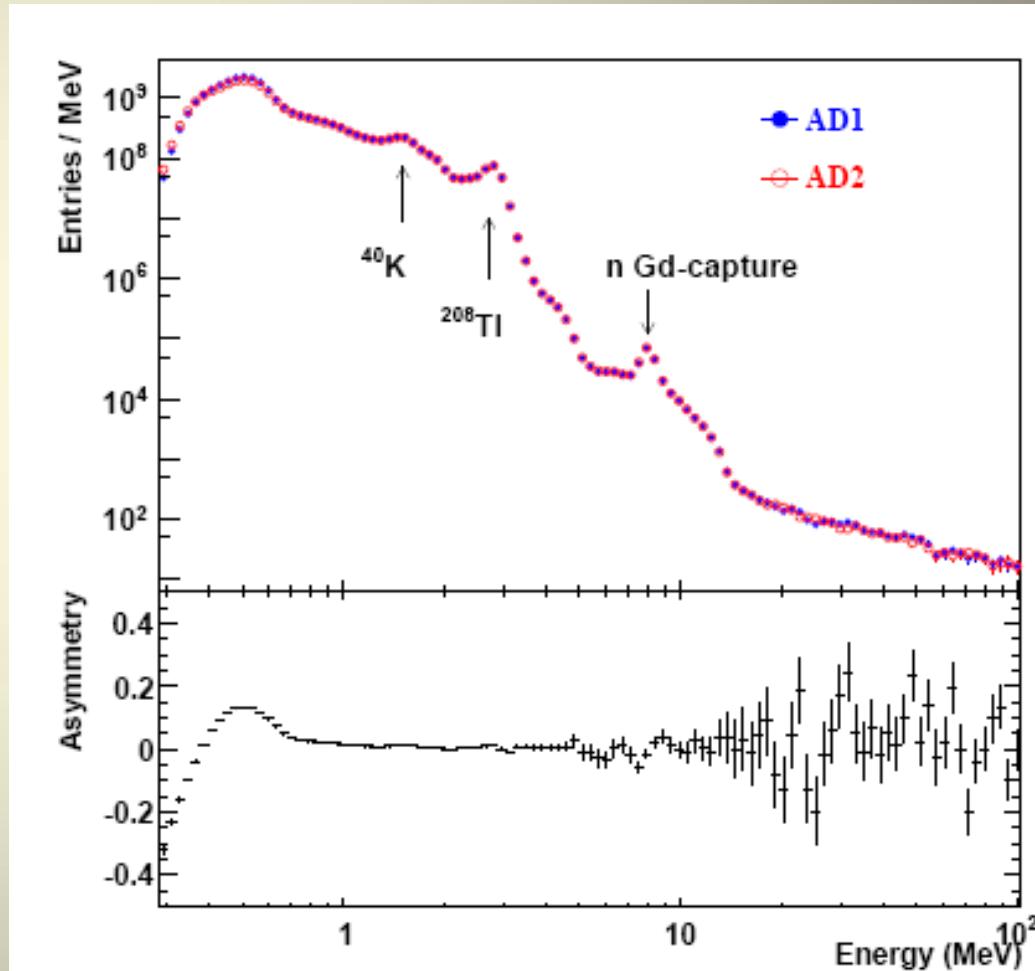
Triggered signals dominated by low-energy radioactivity

Measured Rates:

~65 Hz in each detector
(>0.7 MeV)

Sources:

Stainless Steel: U/Th chains
PMTs: ^{40}K , U/Th chains
Scintillator: Radon/U/Th chains



Antineutrino (IBD) Selection

Use IBD Prompt + Delayed correlated signal to select antineutrinos

Selection:

- Reject Flashers
- Prompt Positron: $0.7 \text{ MeV} < E_p < 12 \text{ MeV}$
- Delayed Neutron: $6.0 \text{ MeV} < E_d < 12 \text{ MeV}$
- Capture time: $1 \mu\text{s} < \Delta t < 200 \mu\text{s}$
- Muon Veto:

Pool Muon: Reject 0.6ms

AD Muon (>20 MeV): Reject 1ms

AD Shower Muon (>2.5GeV): Reject 1s

- Multiplicity:

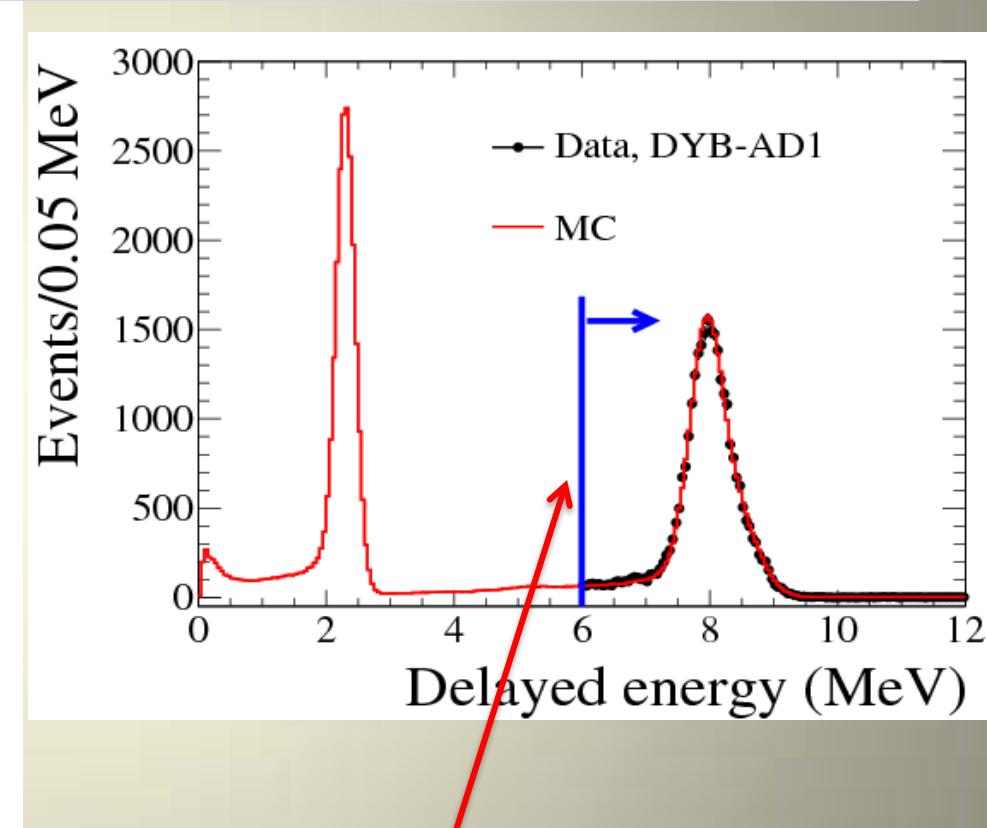
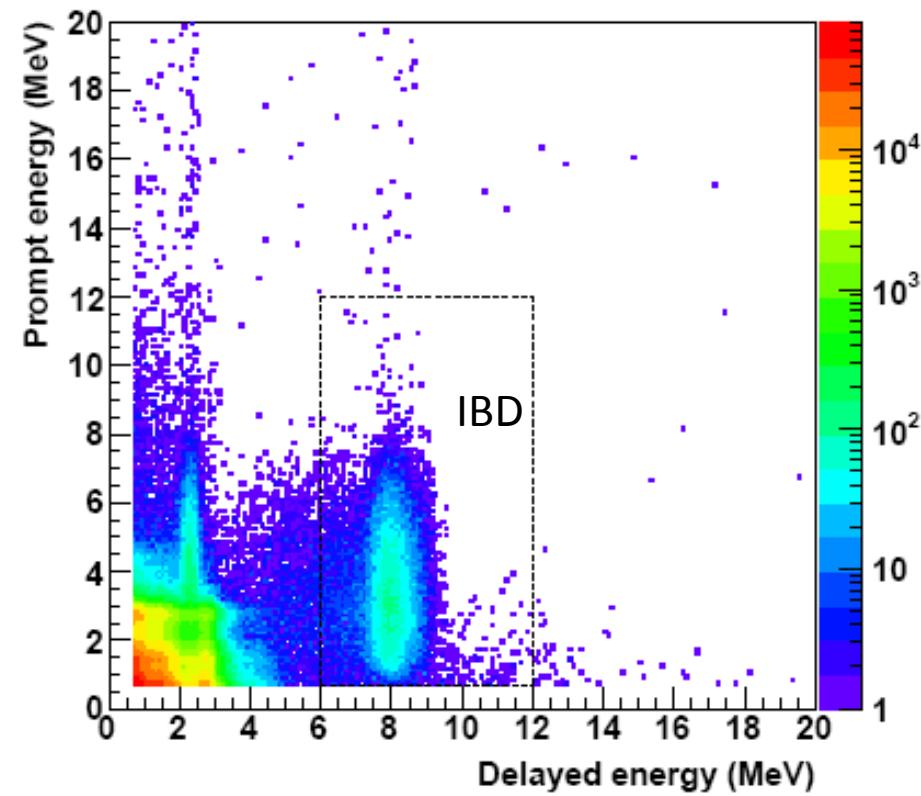
No other signal $> 0.7 \text{ MeV}$
in $-200 \mu\text{s}$ to $200 \mu\text{s}$ of IBD.

Selection driven
by uncertainty in
relative detector
efficiency

$$\frac{N_f}{N_n} = \left(\frac{N_{p,f}}{N_{p,n}} \right) \left(\frac{L_n}{L_f} \right)^2 \left(\frac{\epsilon_f}{\epsilon_n} \right) \left[\frac{P_{\text{sur}}(E, L_f)}{P_{\text{sur}}(E, L_n)} \right]$$

Prompt/Delayed Energy

Clear separation of antineutrino events from most other signals



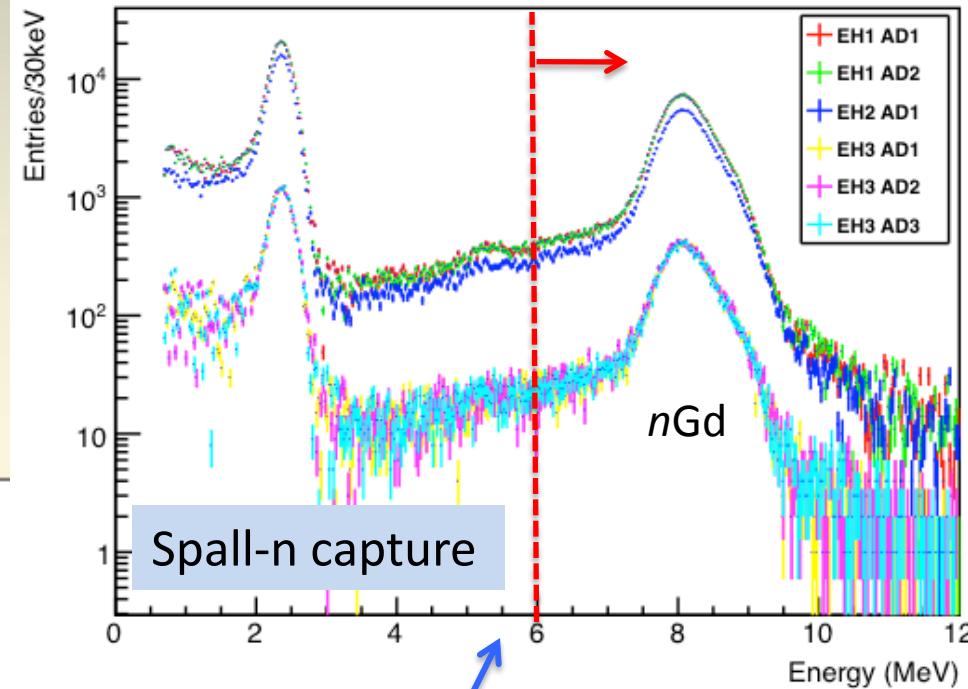
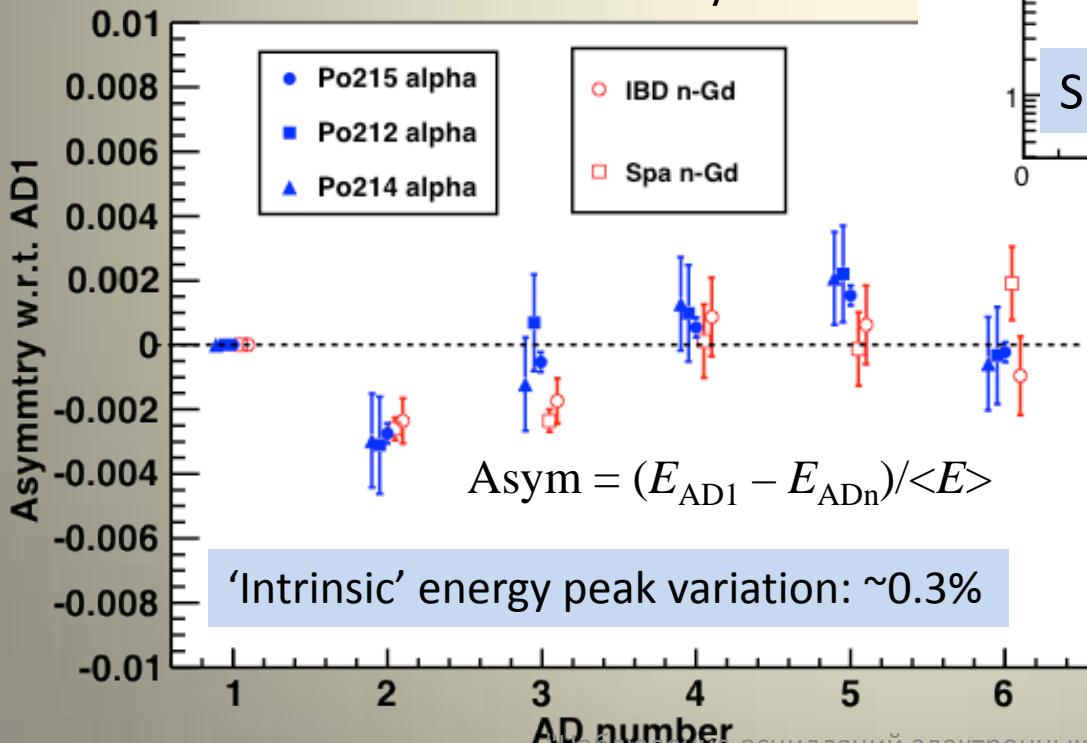
Uncertainty in relative E_d efficiency (0.12%) between detectors is largest systematic.

Delayed Energy Cut

Largest uncertainty between detectors

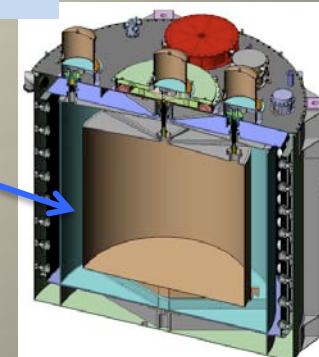
Some n Gd gammas escape scintillator region, visible as tail of n Gd energy peak.

Use variations in energy peaks to constrain relative efficiency.



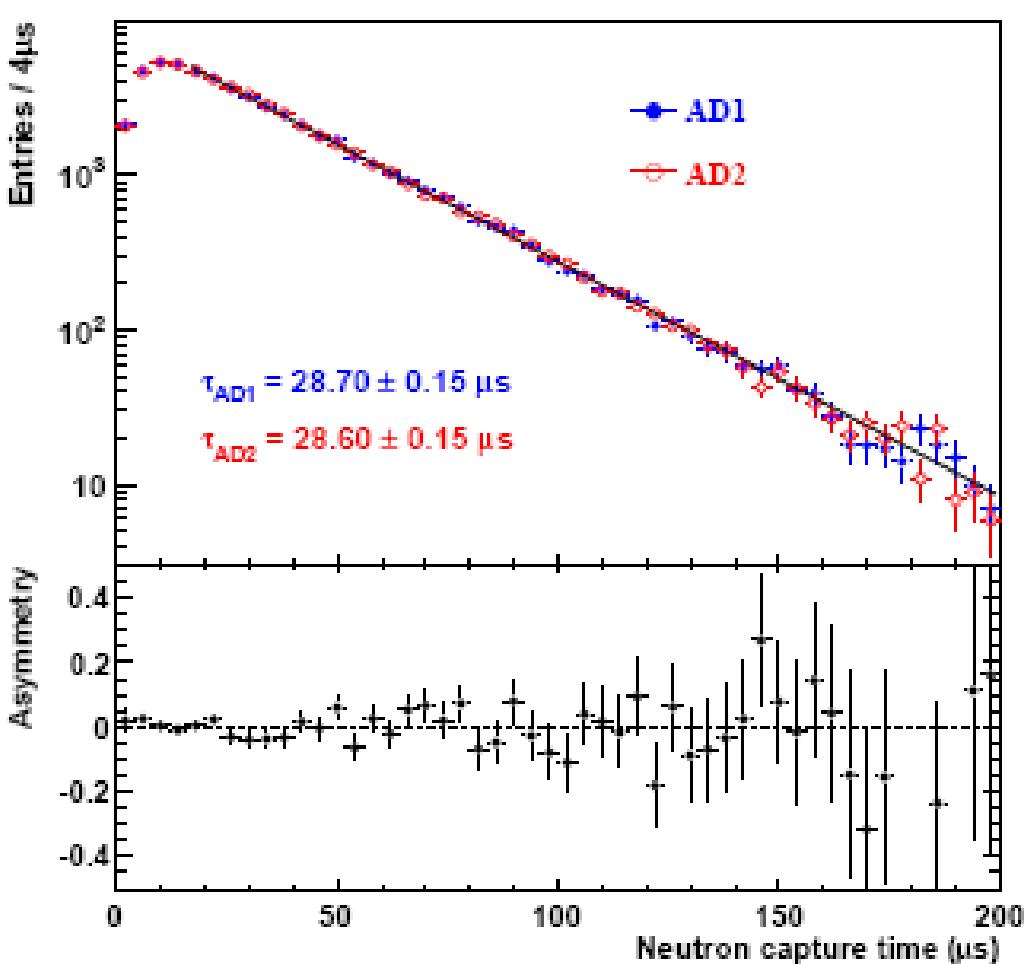
Efficiency variations
estimated at 0.12%

Motivation for
3-zone design



Gd Capture Ratio

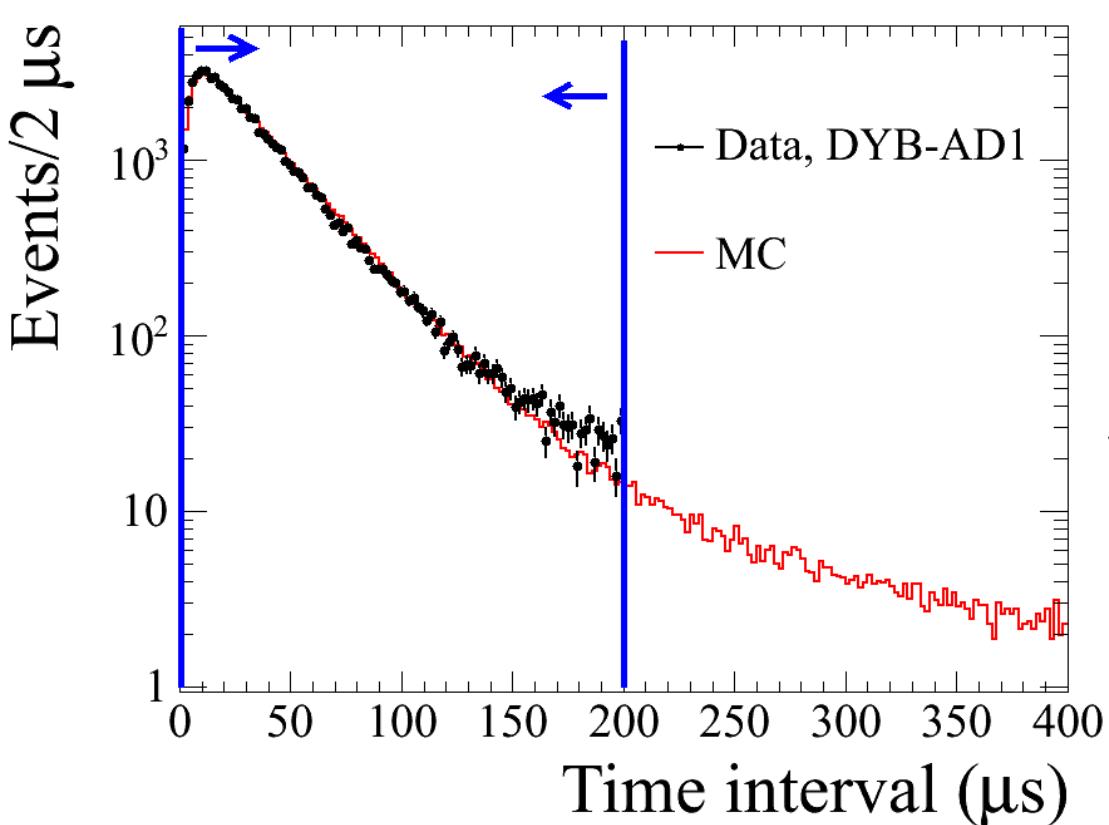
Capture time in each detector constrains H/Gd capture ratio



Measurement of Am-C source neutron capture time distributions constrain uncertainty in relative H/Gd capture efficiency to < 0.1% between detectors.

Capture Time

Consistent IBD capture time measured in all detectors



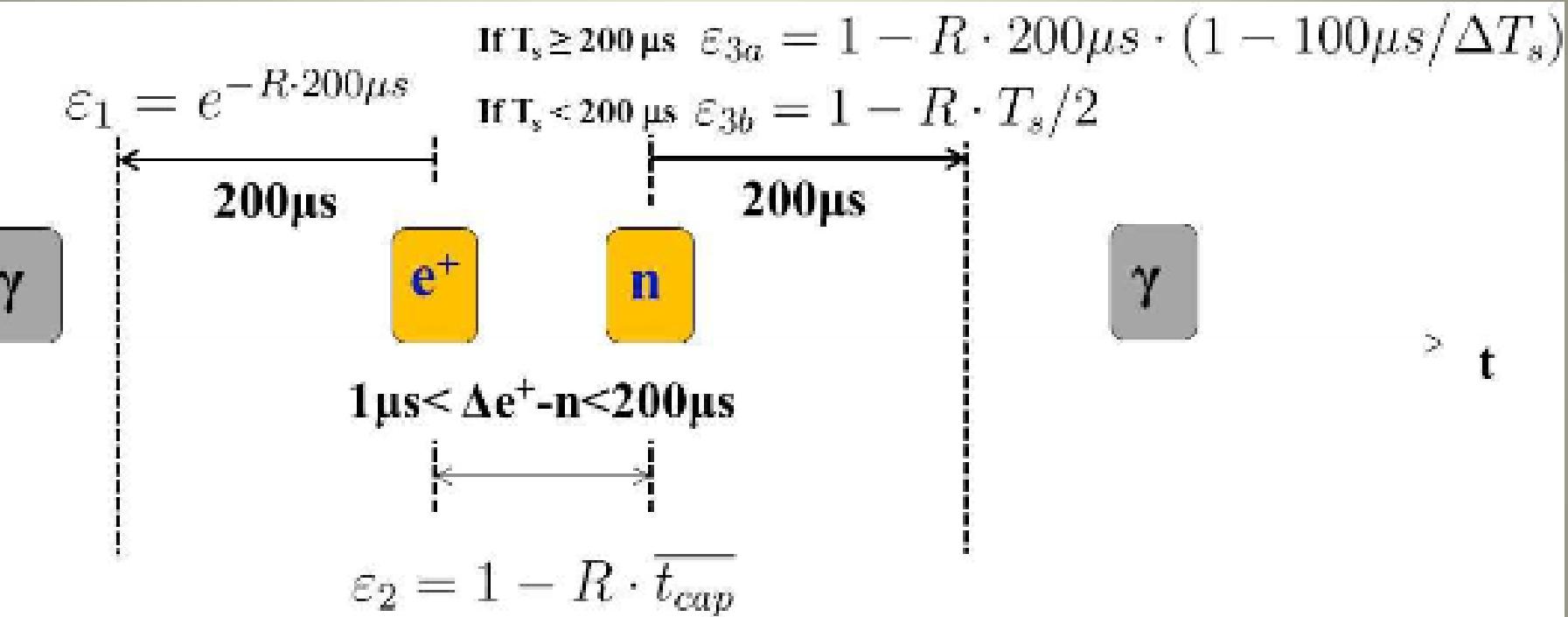
Relative detector efficiency estimated within 0.01% by considering possible variations in Gd concentration.

*Simulation contains no background
(deviates from data at >150 μs)*

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Multiplicity

Ensure exactly one prompt-delayed coincidence



Uncorrelated background and IBD signals result in ambiguous prompt,delayed signals.

-> Reject all IBD with >2 triggers above 0.7 MeV in $-200\mu s$ to $+200\mu s$.

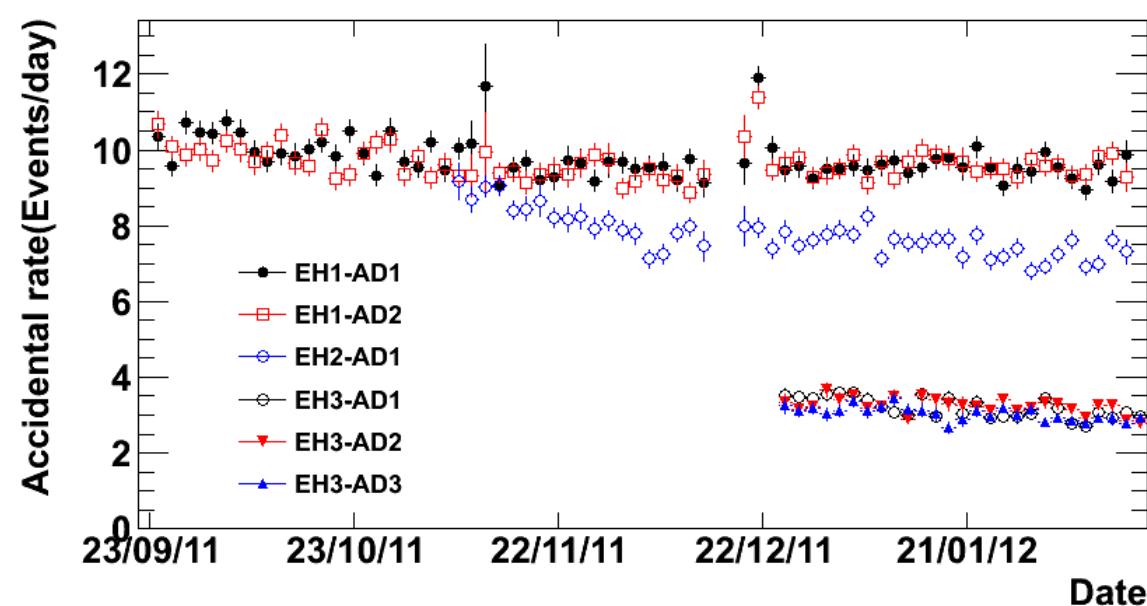
Introduces $\sim 2.5\%$ IBD inefficiency, with negligible uncertainty

Background: Accidentals

Two single signals can accidentally mimic an antineutrino (IBD) signal

Rate and spectrum can be accurately predicted from singles data.

Multiple analyses/methods estimate consistent rates.

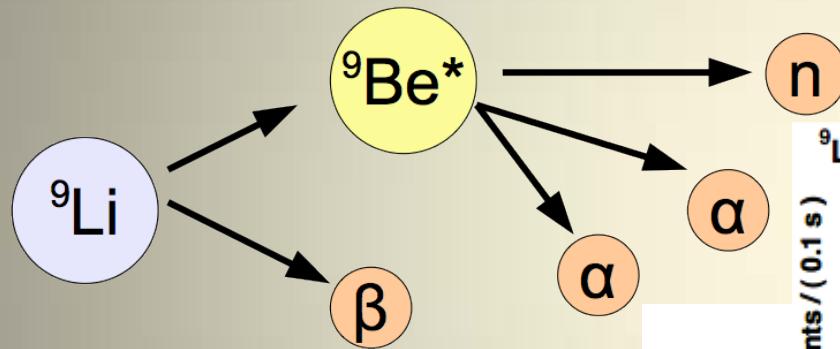


	EH1-AD1	EH1-AD2	EH2-AD1	EH3-AD1	EH3-AD2	EH3-AD3
Accidental rate(/day)	9.82 ± 0.06	9.88 ± 0.06	7.67 ± 0.05	3.29 ± 0.03	3.33 ± 0.03	3.12 ± 0.03
B/S	1.37%	1.38%	1.44%	4.58%	4.77%	4.43%

Background: β -n decay

β -n decay:

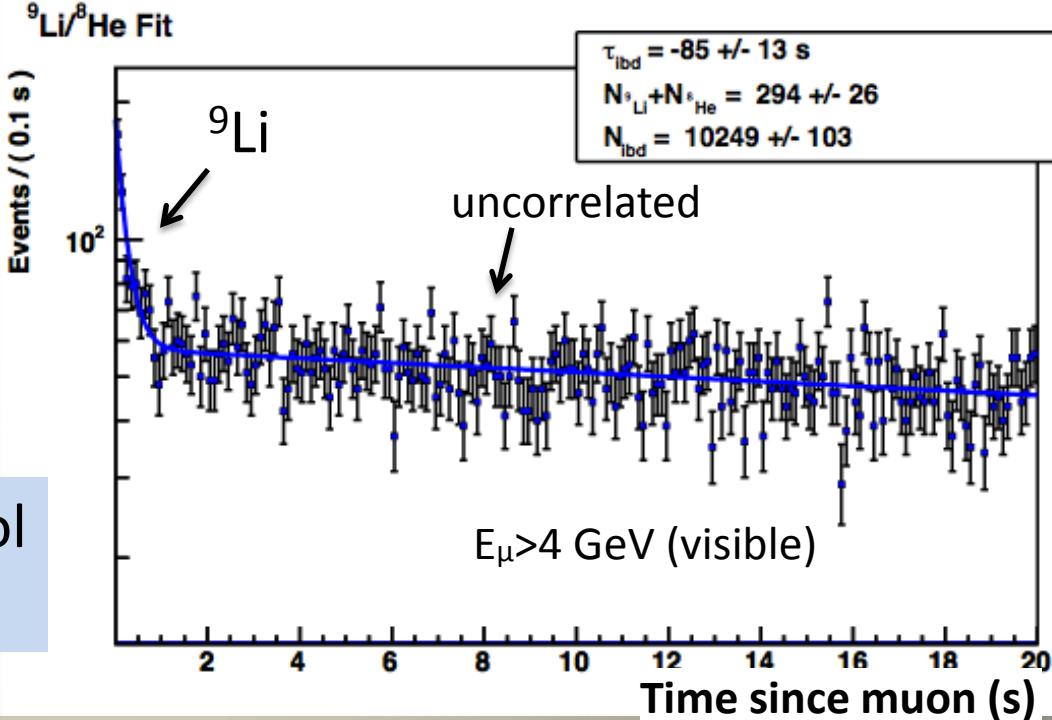
- Prompt: β -decay
- Delayed: neutron capture



${}^9\text{Li}$: $\tau_{1/2} = 178 \text{ ms}$, $Q = 13.6 \text{ MeV}$
 ${}^8\text{He}$: $\tau_{1/2} = 119 \text{ ms}$, $Q = 10.6 \text{ MeV}$

Analysis muon veto cuts control
B/S to $\sim 0.4 \pm 0.2\%$.

- Generated by cosmic rays
- Long-lived
- Mimic antineutrino signal



Background: Fast neutrons

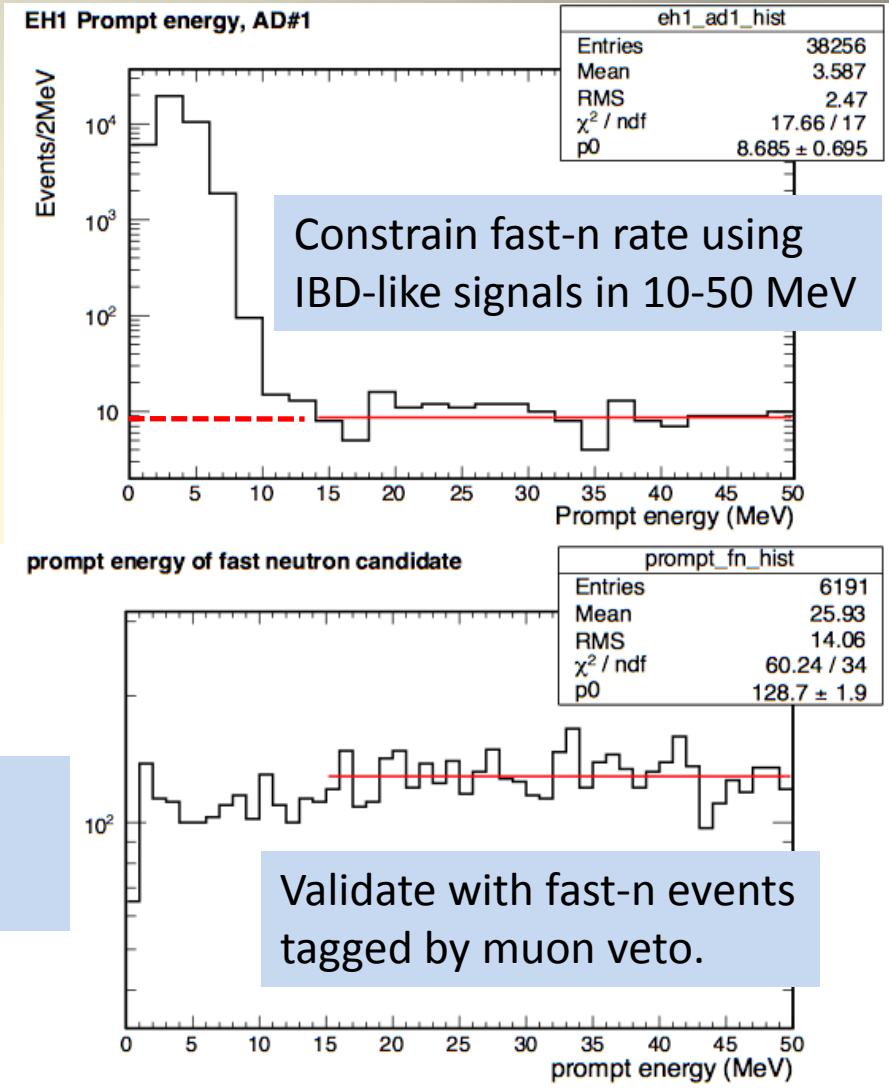
Fast Neutrons:

Energetic neutrons produced by cosmic rays
(inside and outside of muon veto system)

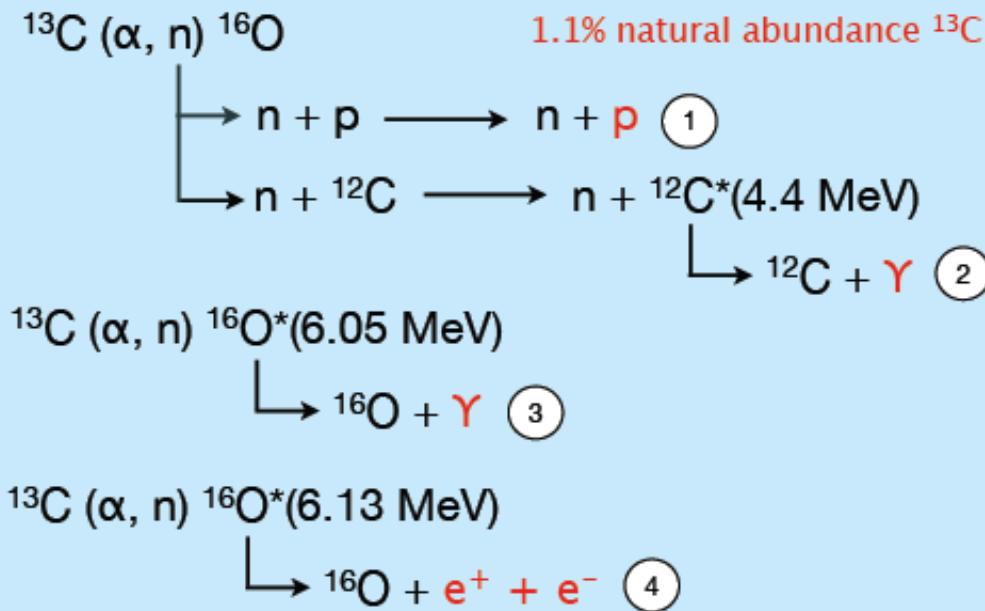
Mimics antineutrino (IBD) signal:

- Prompt: Neutron collides/stops in target
- Delayed: Neutron captures on Gd

Analysis muon veto cuts control B/S
to 0.06% (0.1%) of far (near) signal.



Background: $^{13}\text{C}(\alpha, \text{n})^{16}\text{O}$



Example alpha rate in AD1	^{238}U	^{232}Th	^{235}U	^{210}Po
Bq	0.05	1.2	1.4	10

Potential alpha source:

^{238}U , ^{232}Th , ^{235}U , ^{210}Po :

Each of them are measured in-situ:

U&Th: cascading decay of

Bi(or Rn) – Po – Pb

^{210}Po : spectrum fitting

Combining (α, n) cross-section, correlated background rate is determined.

Near Site: 0.04+-0.02 per day,

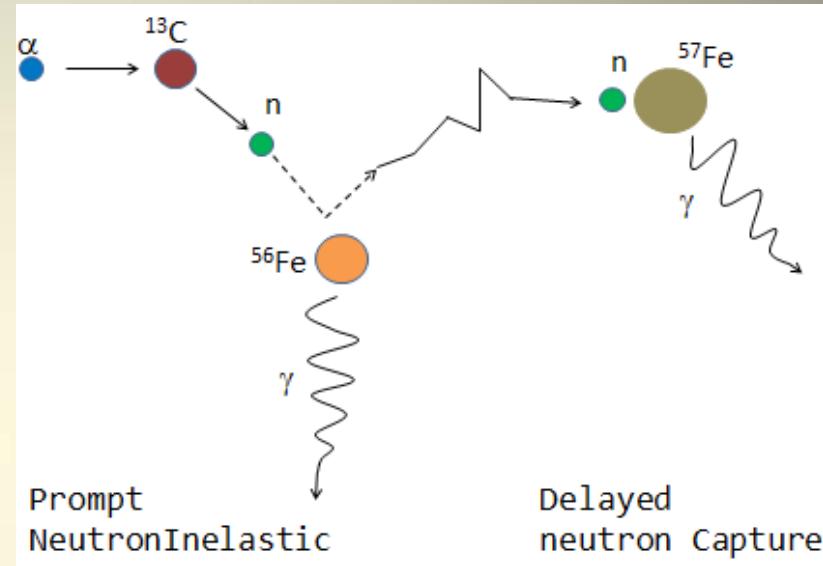
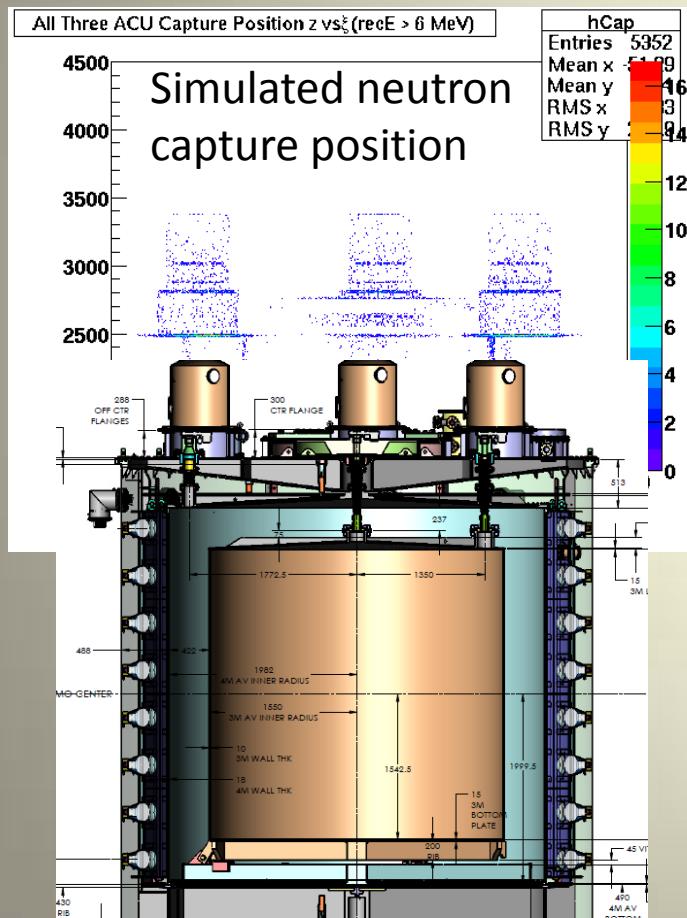
B/S (0.006 ± 0.004)%

Far Site: 0.03+-0.02 per day,

B/S (0.04 ± 0.02)%

Background: ^{241}Am - ^{13}C neutrons

Weak (0.5Hz) neutron source in ACU can mimic IBD via inelastic scattering and capture on iron.



Constrain far site B/S to $0.3 \pm 0.3\%$:

- Measure uncorrelated gamma rays from ACU in data
- Estimate ratio of correlated/uncorrelated rate using simulation
- Assume 100% uncertainty from simulation

Data Set Summary

	AD1	AD2	AD3	AD4	AD5	AD6
Antineutrino candidates	28935	28975	22466	3528	3436	3452
DAQ live time (day)	49.5530		49.4971		48.9473	
Efficiency	0.8019	0.7989	0.8363	0.9547	0.9543	0.9538
Accidentals (/day)	9.82 ± 0.06	9.88 ± 0.06	7.67 ± 0.05	3.29 ± 0.03	3.33 ± 0.03	3.12 ± 0.03
Fast neutron (/day)	0.84 ± 0.28	0.84 ± 0.28	0.74 ± 0.44	0.04 ± 0.04	0.04 ± 0.04	0.04 ± 0.04
$^8\text{He}/^9\text{Li}$ (/day)	3.1 ± 1.6		1.8 ± 1.1	0.16 ± 0.11		
Am-C corr. (/day)	0.2 ± 0.2					
$^{13}\text{C}(\alpha, n)^{16}\text{O}$ (/day)	0.04 ± 0.02	0.04 ± 0.02	0.035 ± 0.02	0.03 ± 0.02	0.03 ± 0.02	0.03 ± 0.02
Antineutrino rate (/day)	714.17 ± 4.58	717.86 ± 4.60	532.29 ± 3.82	71.78 ± 1.29	69.80 ± 1.28	70.39 ± 1.28

Consistent rates for side-by-side detectors

Uncertainty currently dominated by statistics

Reactor Flux Expectation

Antineutrino flux is estimated for each reactor core

Flux estimated using:

$$S(E_\nu) = \frac{W_{th}}{\sum_i (f_i/F)e_i} \sum_i^{istopes} (f_i/F) S_i(E_\nu)$$

Reactor operators provide:

- Thermal power data: W_{th}
- Relative isotope fission fractions: f_i

Energy released per fission: e_i

V. Kopekin et al., Phys. Atom. Nucl. 67, 1892 (2004)

Antineutrino spectra per fission: $S_i(E_\nu)$

K. Schreckenbach et al., Phys. Lett. B160, 325 (1985)

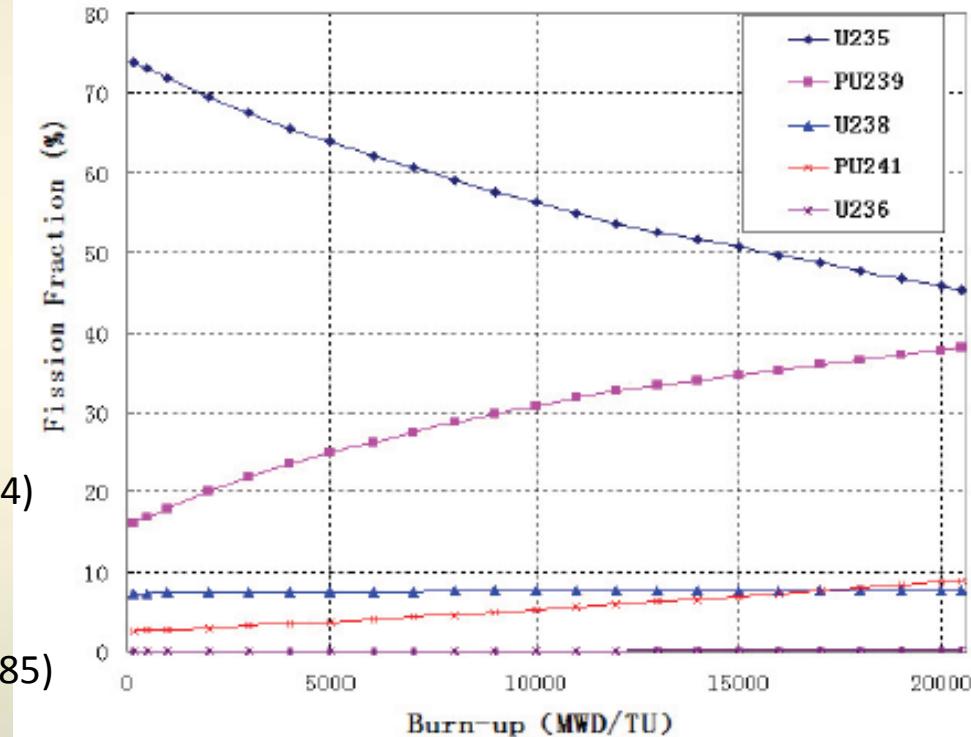
A. A. Hahn et al., Phys. Lett. B218, 365 (1989)

P. Vogel et al., Phys. Rev. C24, 1543 (1981)

T. Mueller et al., Phys. Rev. C83, 054615 (2011)

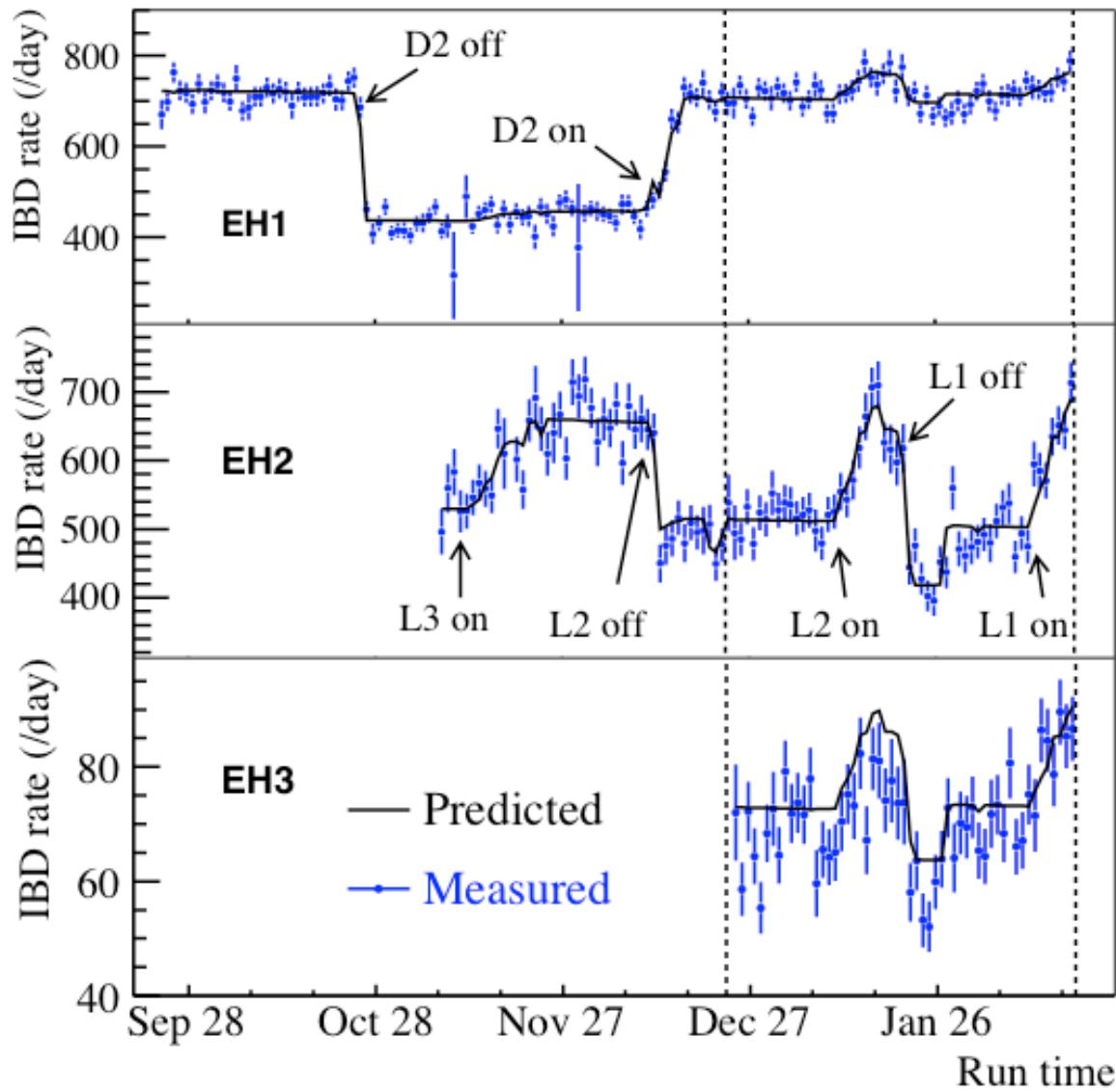
P. Huber, Phys. Rev. C84, 024617 (2011)

Isotope fission rates vs. reactor burnup



Flux model has negligible impact on far vs. near oscillation measurement

Antineutrino Rate vs. Time

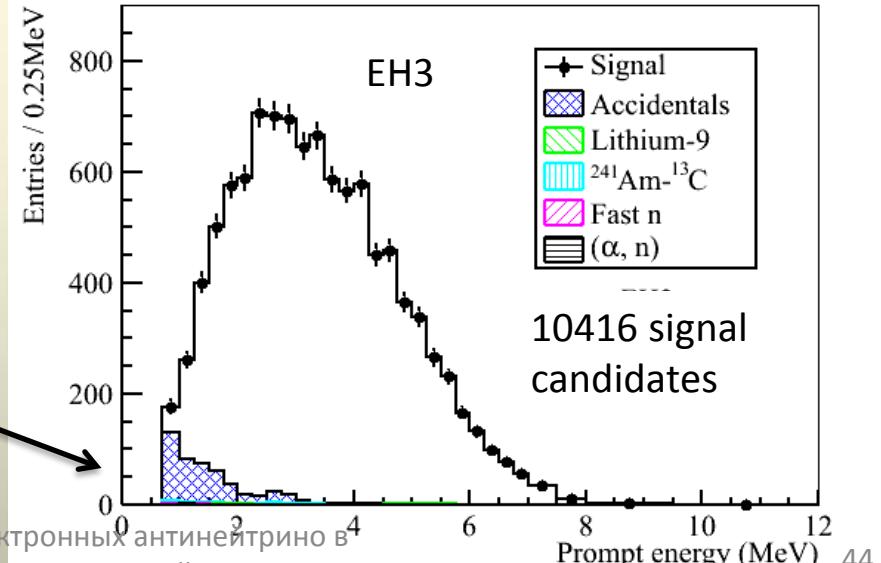
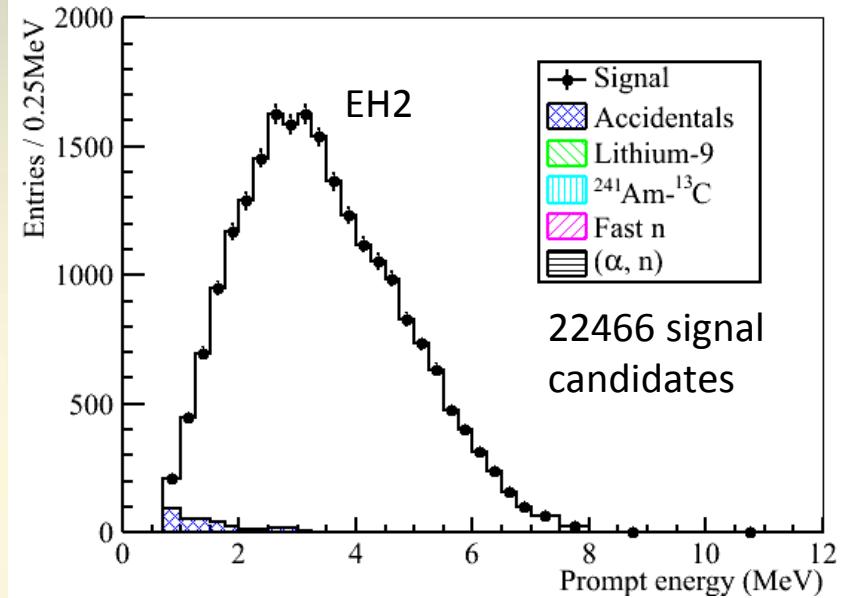
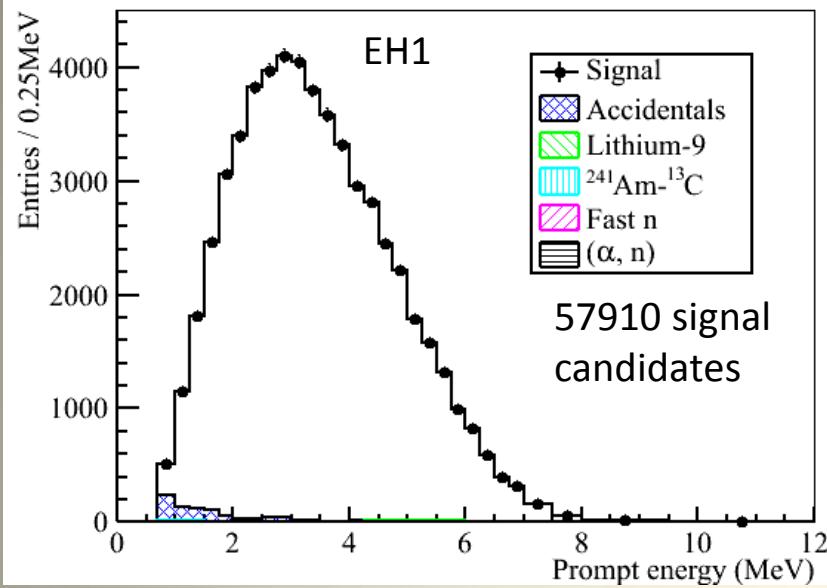


Detected rate strongly correlated with reactor flux expectations.

Predicted Rate: (in figure)

- Assumes no oscillation.
- Normalization is determined by fit to data.
- Absolute normalization is within a few percent of expectations.

Prompt Positron Spectra



High-statistics reactor antineutrino spectra.

B/S ratio is 5% (2%) at far (near) sites.

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Uncertainty Summary

Detector			
	Efficiency	Correlated	Uncorrelated
Target Protons		0.47%	0.03%
Flasher cut	99.98%	0.01%	0.01%
Delayed energy cut	90.9%	0.6%	0.12%
Prompt energy cut	99.88%	0.10%	0.01%
Multiplicity cut		0.02%	<0.01%
Capture time cut	98.6%	0.12%	0.01%
Gd capture ratio	83.8%	0.8%	<0.1%
Spill-in	105.0%	1.5%	0.02%
Livetime	100.0%	0.002%	<0.01%
Combined	78.8%	1.9%	0.2%

For near/far oscillation, only uncorrelated uncertainties are used.

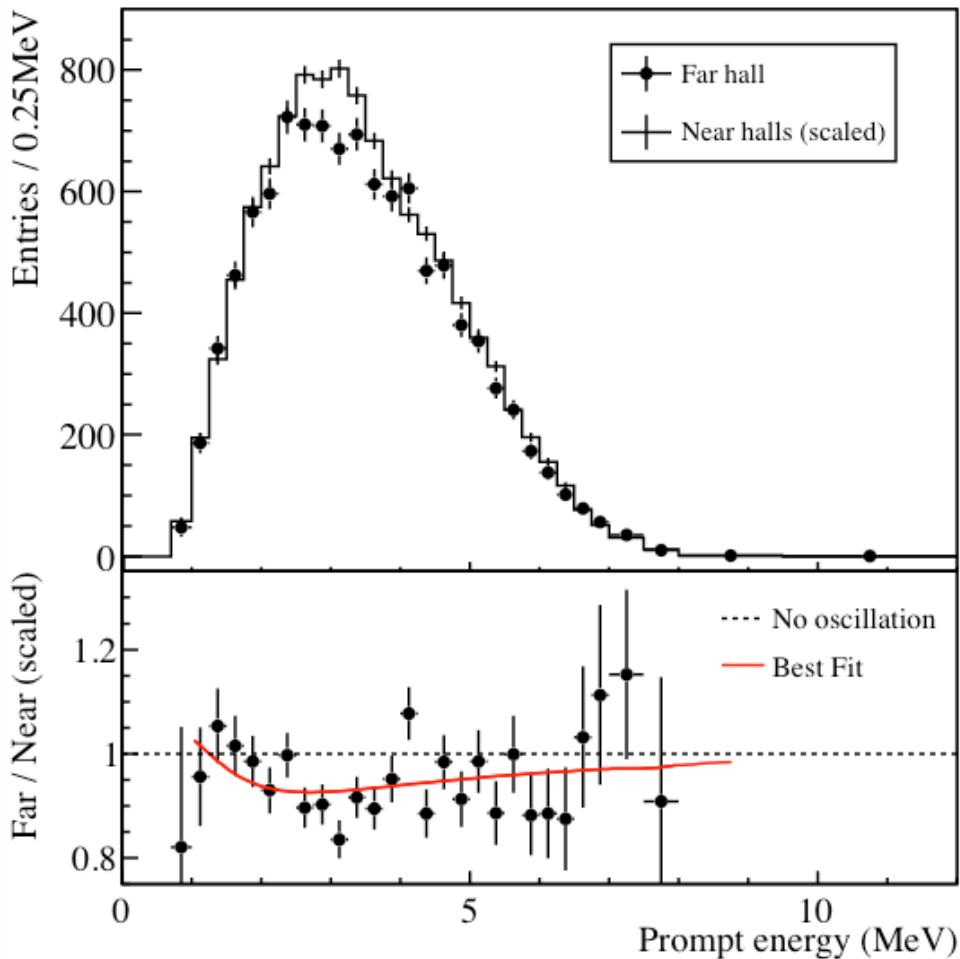
Reactor		
Correlated		Uncorrelated
Energy/fission	0.2%	Power 0.5%
$\bar{\nu}_e$ /fission	3%	Fission fraction 0.6%
		Spent fuel 0.3%
Combined	3%	Combined 0.8%

Largest systematics are smaller than far site statistics (~1%)

Influence of uncorrelated reactor systematics reduced by far vs. near measurement.

Far vs. Near Comparison

Compare the far/near measured rates and spectra



$$R = \frac{Far_{measured}}{Far_{expected}} = \frac{M_4 + M_5 + M_6}{\sum_{i=4}^6 (\alpha_i(M_1 + M_2) + \beta_i M_3)}$$

M_n are the measured rates in each detector.
 Weights α_i, β_i are determined from baselines
 and reactor fluxes.

$$R = 0.940 \pm 0.011 \text{ (stat)} \pm 0.004 \text{ (syst)}$$

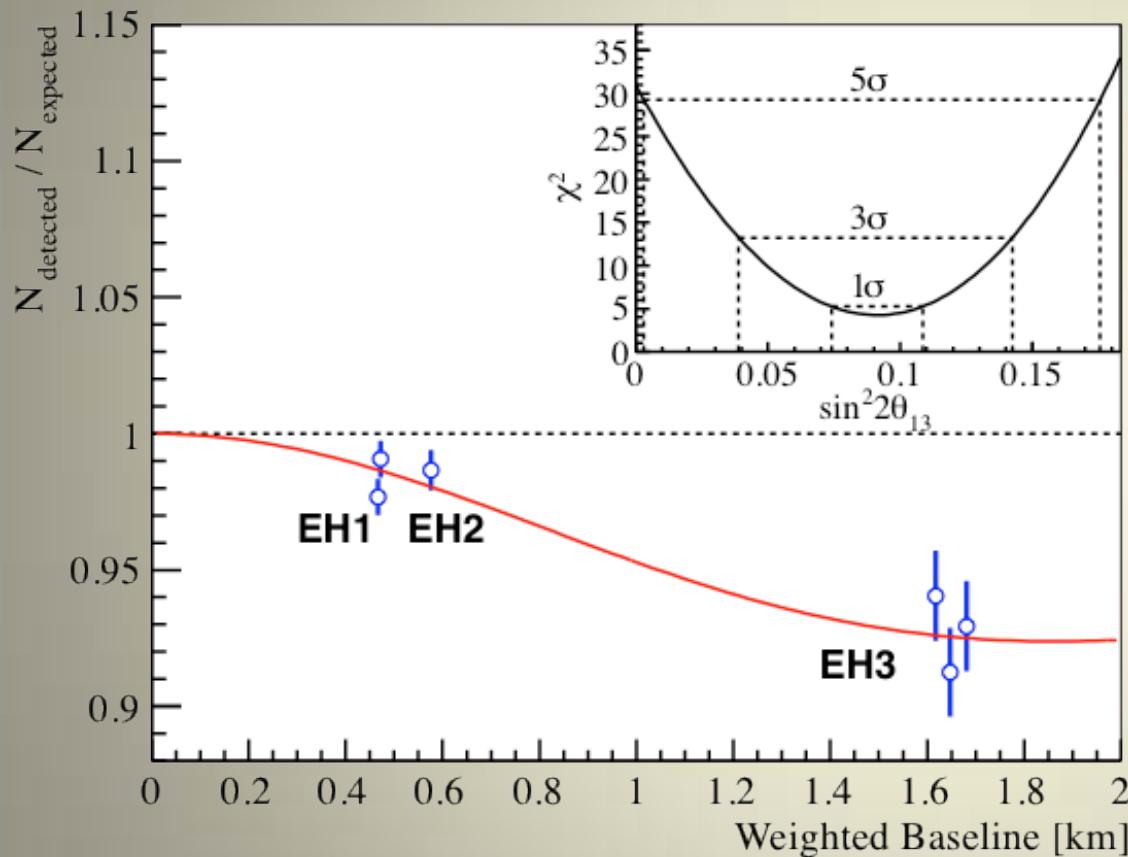
Clear observation of far site deficit.

Spectral distortion consistent with oscillation.*

* Caveat: Spectral systematics not fully studied;
 θ_{13} value from shape analysis is not recommended.

Rate Analysis

Estimate θ_{13} using measured rates in each detector.



Uses standard χ^2 approach.

Far vs. near relative measurement.
[Absolute rate is not constrained.]

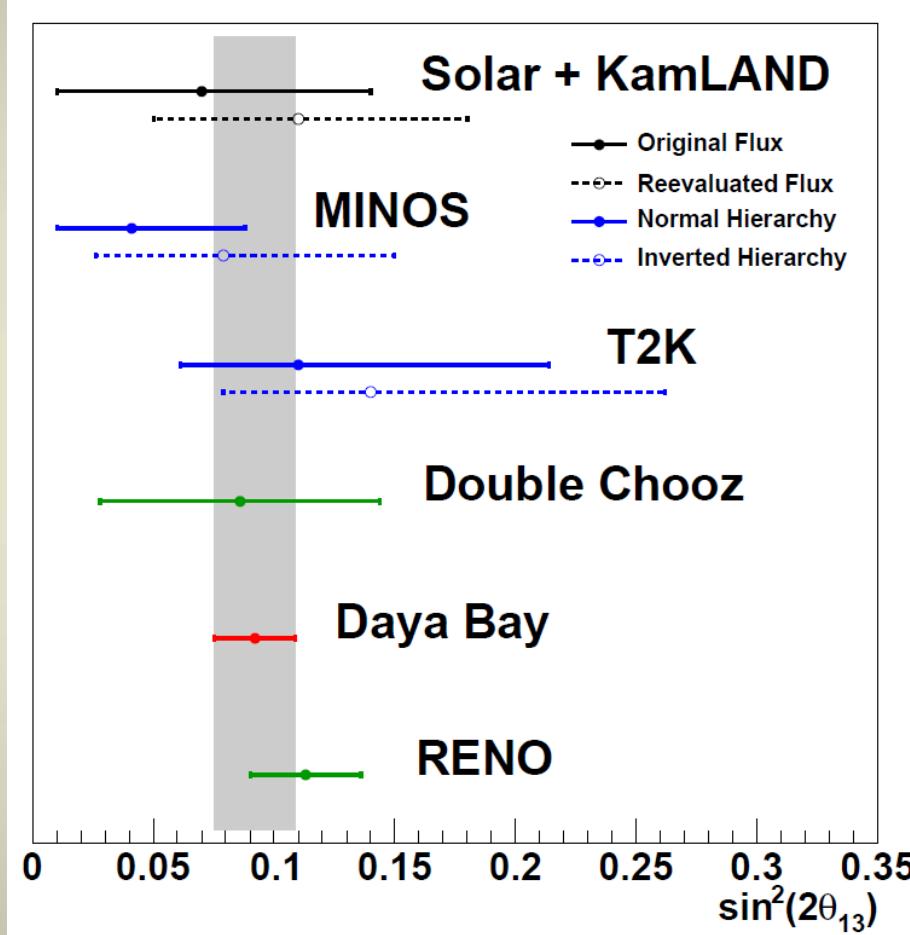
Consistent results obtained by independent analyses, different reactor flux models.

$\sin^2 2\theta_{13} = 0$
excluded at 5.2σ

$$\sin^2 2\theta_{13} = 0.092 \pm 0.016 \text{ (stat)} \pm 0.005 \text{ (syst)}$$

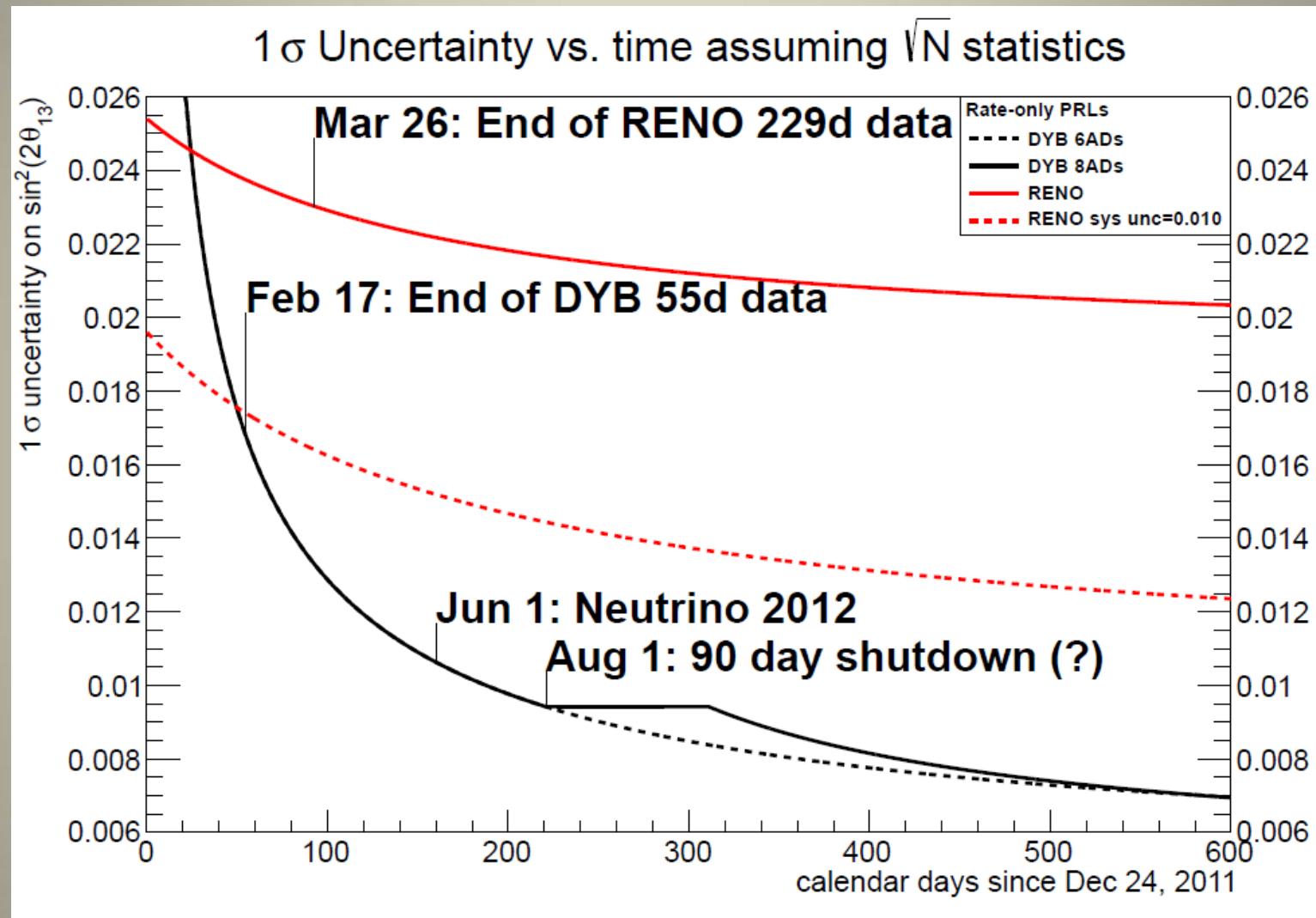
θ_{13}

Daya Bay precision surpasses all existing estimates.



Expect more statistics and further improvements in analysis.

Projected sensitivity to θ_{13}



Summary

- The Daya Bay reactor neutrino experiment has made an unambiguous observation of reactor electron-antineutrino disappearance at ~ 2 km:

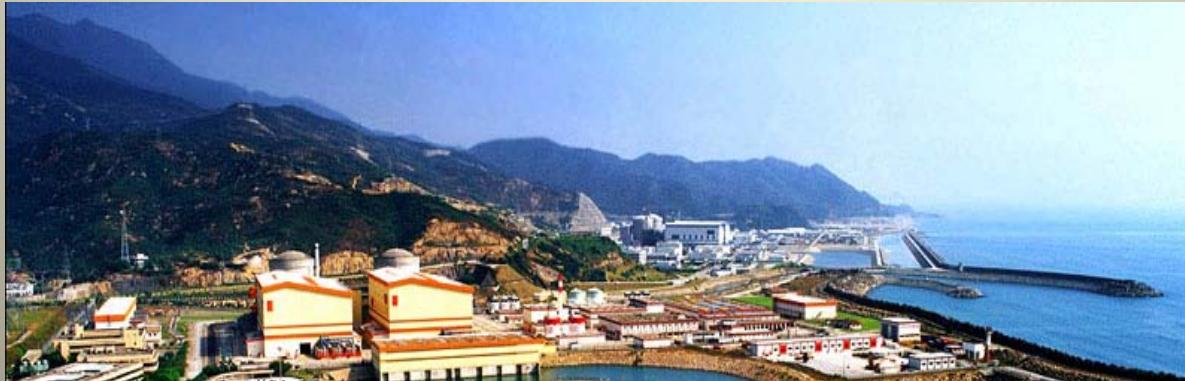
$$R = 0.940 \pm 0.011 \text{ (stat)} \pm 0.004 \text{ (syst)}$$

- Interpretation of disappearance as neutrino oscillation yields:

$$\sin^2 2\theta_{13} = 0.092 \pm 0.016 \text{ (stat)} \pm 0.005 \text{ (syst)}$$

ruling out zero at 5.2 standard deviations.

- Installation of final pair of antineutrino detectors this summer



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